

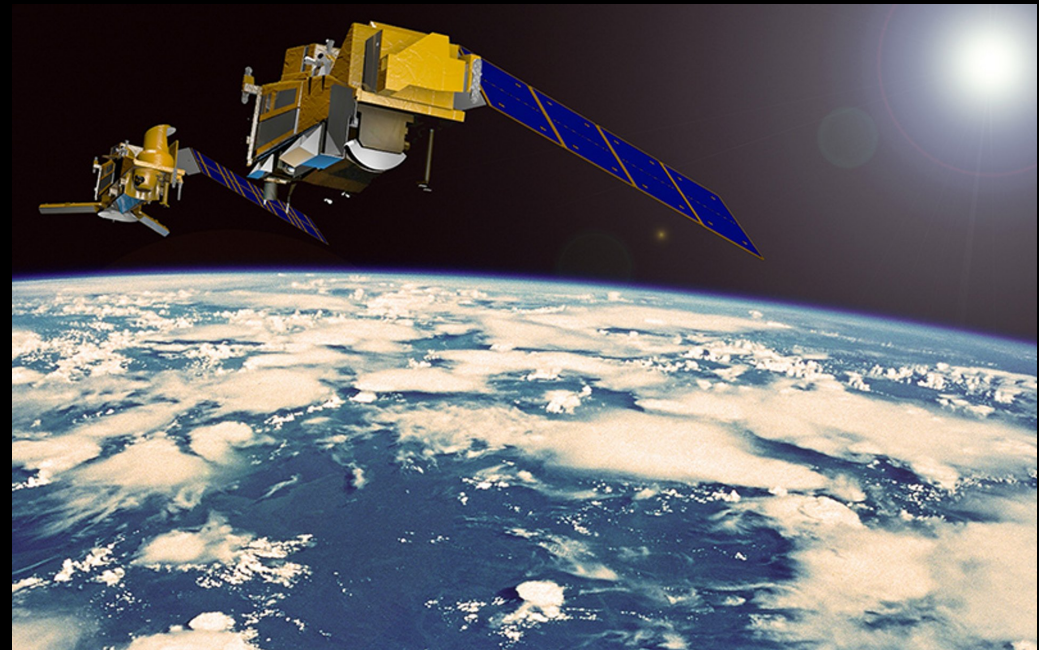


CHALMERS
UNIVERSITY OF TECHNOLOGY

Characterising and reducing uncertainties in all-sky radiative transfer

Vasileios Barlakas

Annual Fellow Day
01.03.2021



CREATE – Characterizing and REducing uncertainties in All-sky microwave radiative TransfEr

● Sensors

- Sensors with frequencies up to ~183 GHz
- Ice Cloud Imager (ICI): 183.31 – 664 GHz

● Applications

- Stand-alone retrievals
- Data assimilation (DA)

● Open issues

- Three-dimensional radiative transfer is ingored
- Hydrometeor orientation and polarization are ignored
- RTTOV-SCATT accuracy
 - Two-stream delta-Eddington approximation
 - Sub-grid variability and cloud overlap scheme

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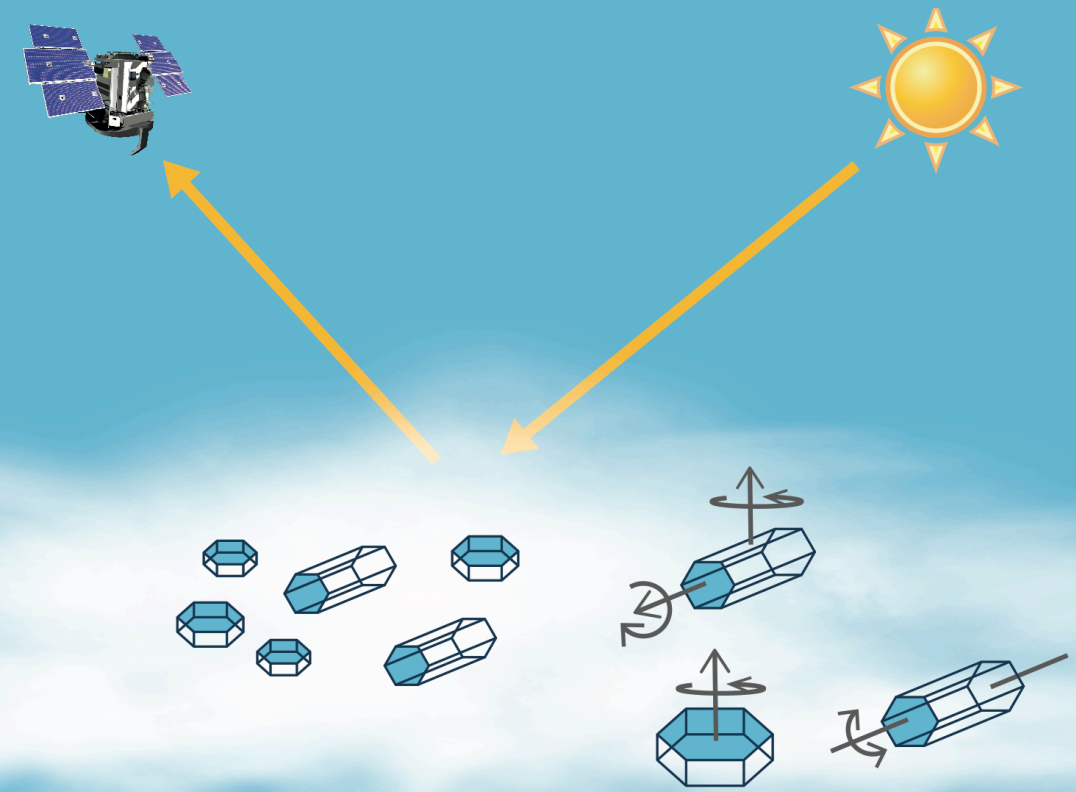
¹Barlakas and Eriksson., *Remote Sens.*, 2020

²Barlakas et al., *AMTD*, 2020

²Geer, Barlakas et al., *to be submitted to GMD*, 2021

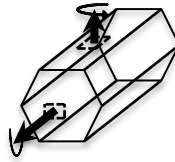
²Barlakas et al., *in preparation, JQSRT*, 2021

Introducing hydrometeor orientation into all-sky millimeter/sub-millimeter assimilation



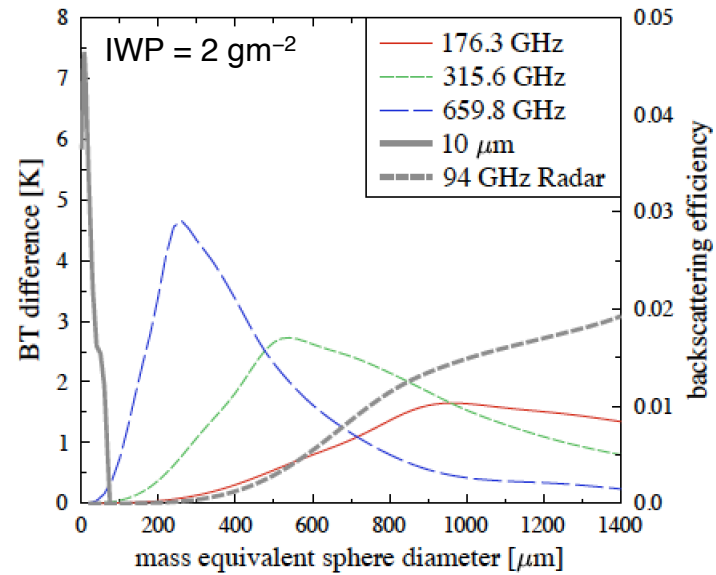
● Why ice clouds?

- Model uncertainties (shape and orientation).



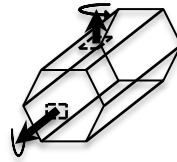
● Why MW and sub-mm?

- Sensitive to large and small ice hydrometeors
- Ice Cloud Imager (ICI)
 - 183–664 GHz + dual polarization
 - Improved ice retrievals + extend data assimilation



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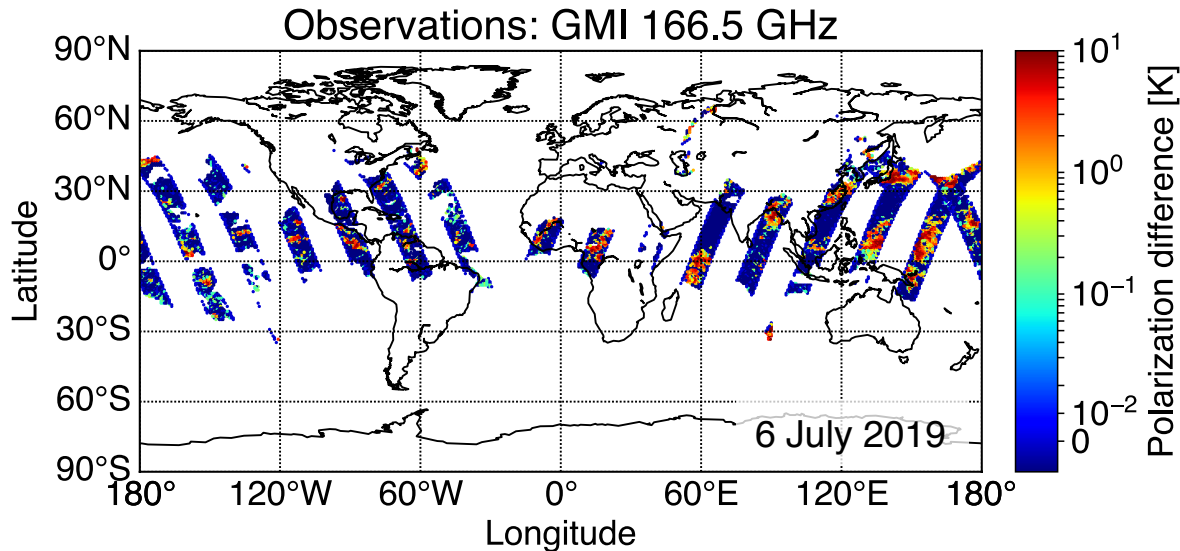
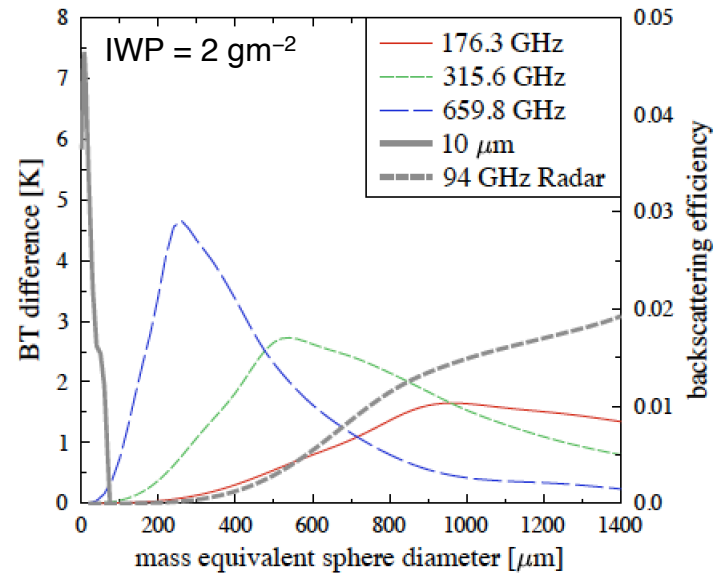
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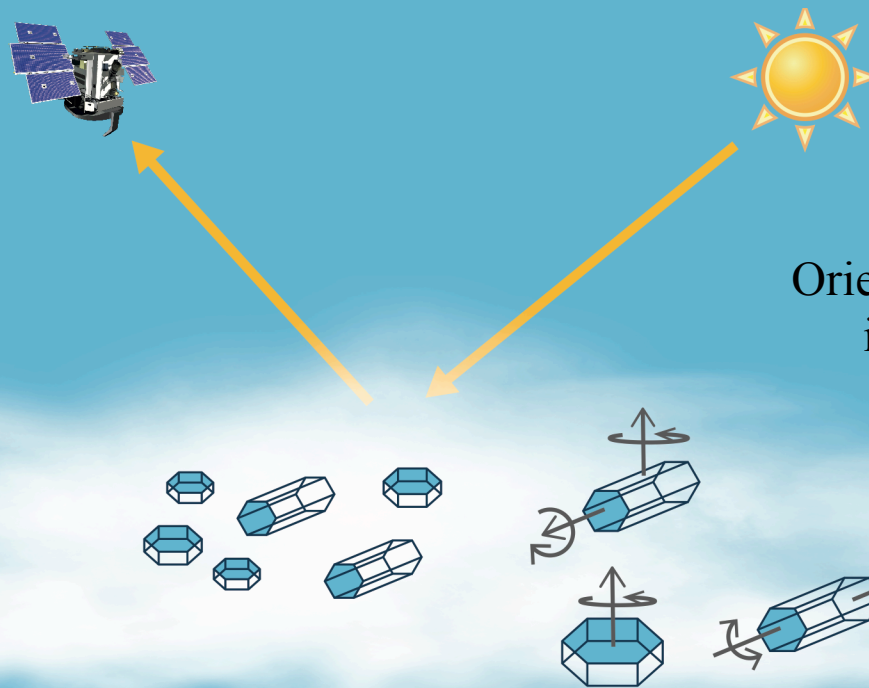
- Sensitive to large and small ice hydrometeors
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+ Polarization



● Assumptions in DA

- Totally randomly oriented (TRO) hydrometeors only
- “Scalar” simulations, i.e., V– or H–polarization
- Invariant scattering properties between V– and H–simulations



Oriented ice hydrometeors induce polarization

$$PD = T_{BV} - T_{BH}$$

- Attenuation driven by the extinction matrix
- ✓ incident direction
- ✓ orientation

$$K = \begin{pmatrix} K_{11} & K_{12} & K_{13} & K_{14} \\ K_{21} & K_{22} & K_{23} & K_{24} \\ K_{31} & K_{32} & K_{33} & K_{34} \\ K_{41} & K_{42} & K_{43} & K_{44} \end{pmatrix}$$

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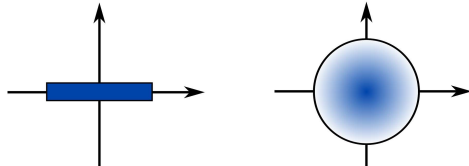
○ Total random orientation (TRO)

- ✗ favored direction
- ✗ orientation

$$K = \begin{pmatrix} K_{11} & 0 & 0 & 0 \\ 0 & K_{11} & 0 & 0 \\ 0 & 0 & K_{11} & 0 \\ 0 & 0 & 0 & K_{11} \end{pmatrix}$$

K_{11} , extinction cross section

Basic orientation



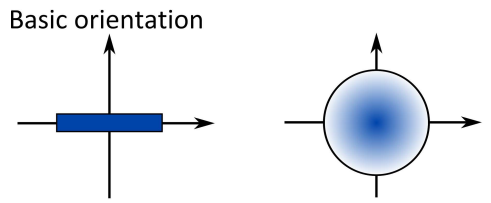
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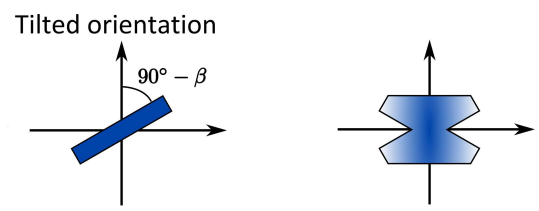
K_{11} , extinction cross section



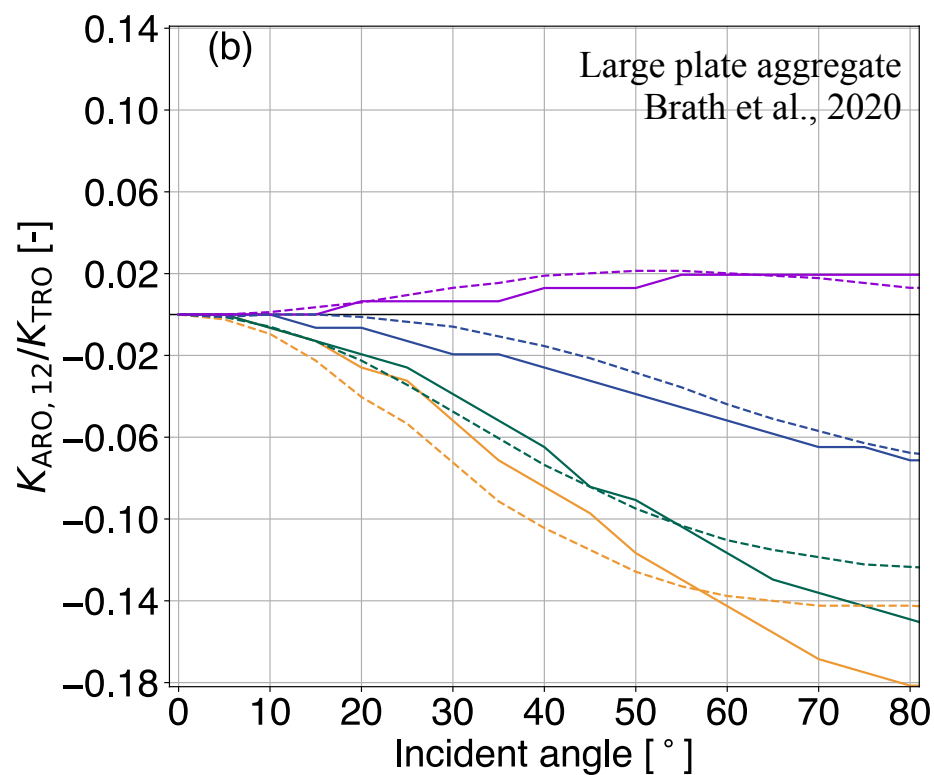
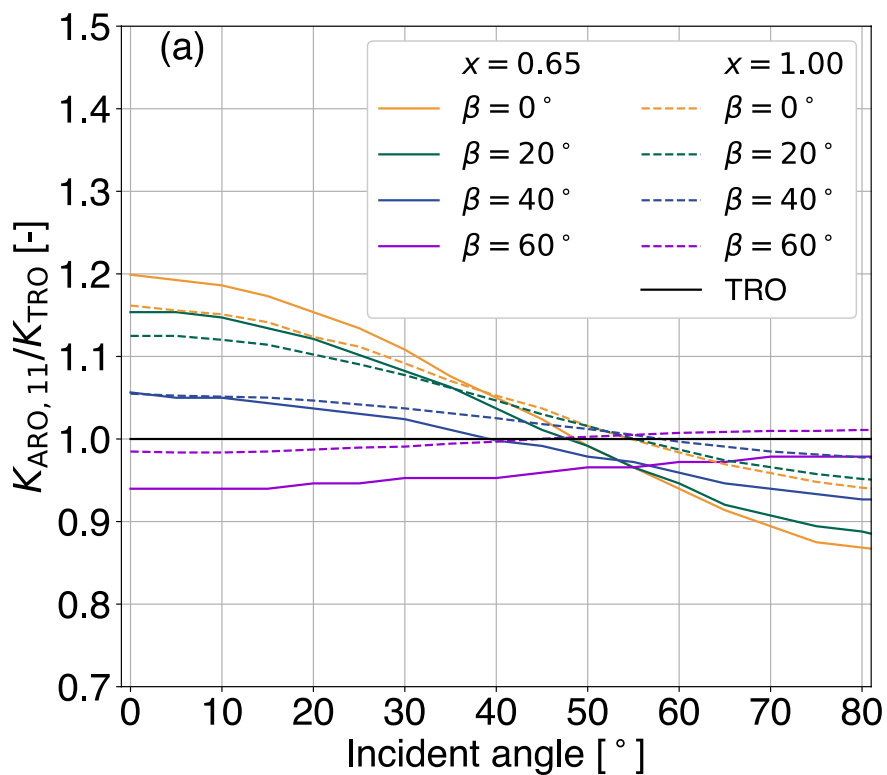
- Azimuthal random orientation (ARO)
- ✓ orientation based on tilt angle β
- ✗ orientation in the azimuth

$$K = \begin{pmatrix} K_{11} & K_{12} & 0 & 0 \\ K_{12} & K_{11} & 0 & 0 \\ 0 & 0 & K_{11} & K_{34} \\ 0 & 0 & -K_{34} & K_{11} \end{pmatrix}$$

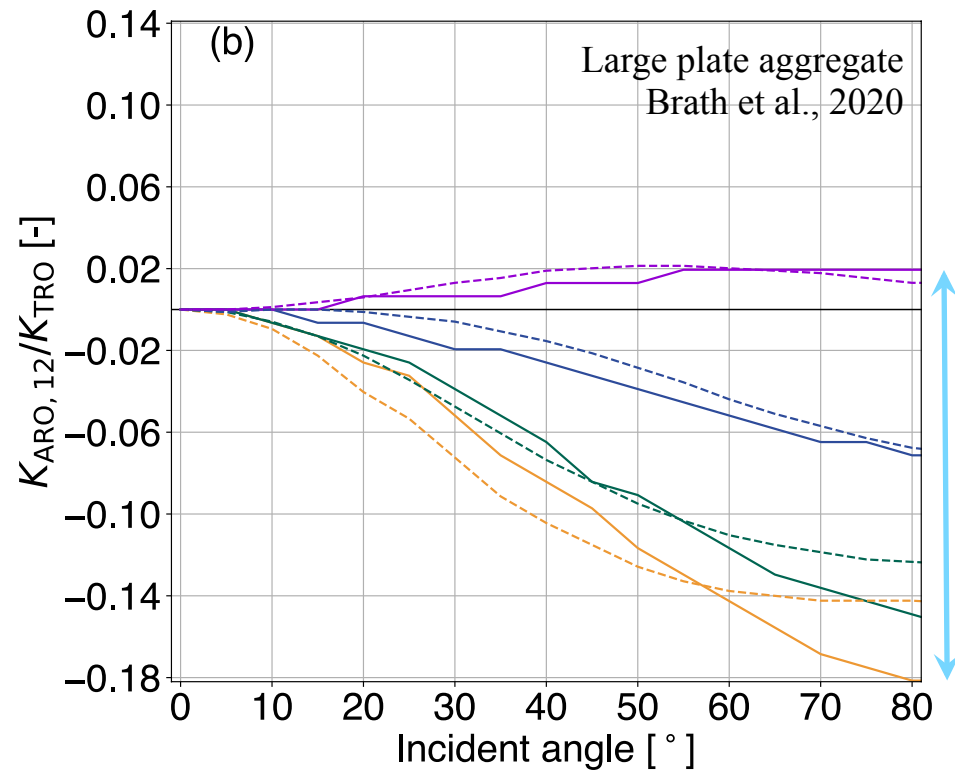
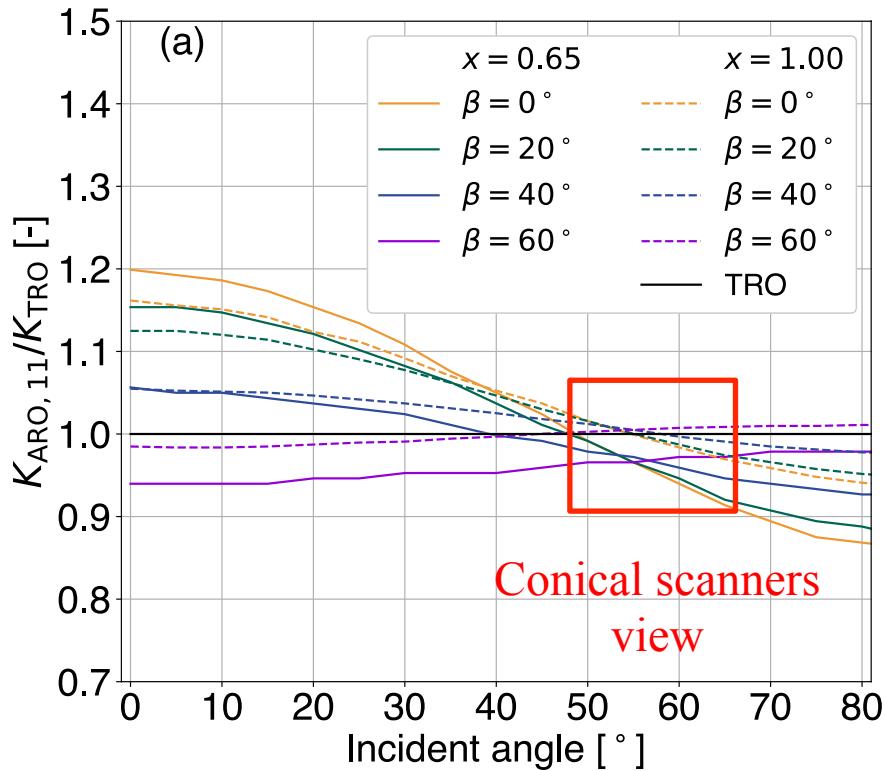
K_{12} , cross section for linear polarisation
 K_{34} , cross section for circular polarisation



● TRO vs ARO



● **TRO vs ARO**



- Extinction cross section, K_{11} \Rightarrow At $\theta_{inc} \sim 55^\circ$, the differences are close to 0.
- Linear polarization cross section, K_{12} \Rightarrow Large differences up to 18 % between V & H

• Orientation approximation

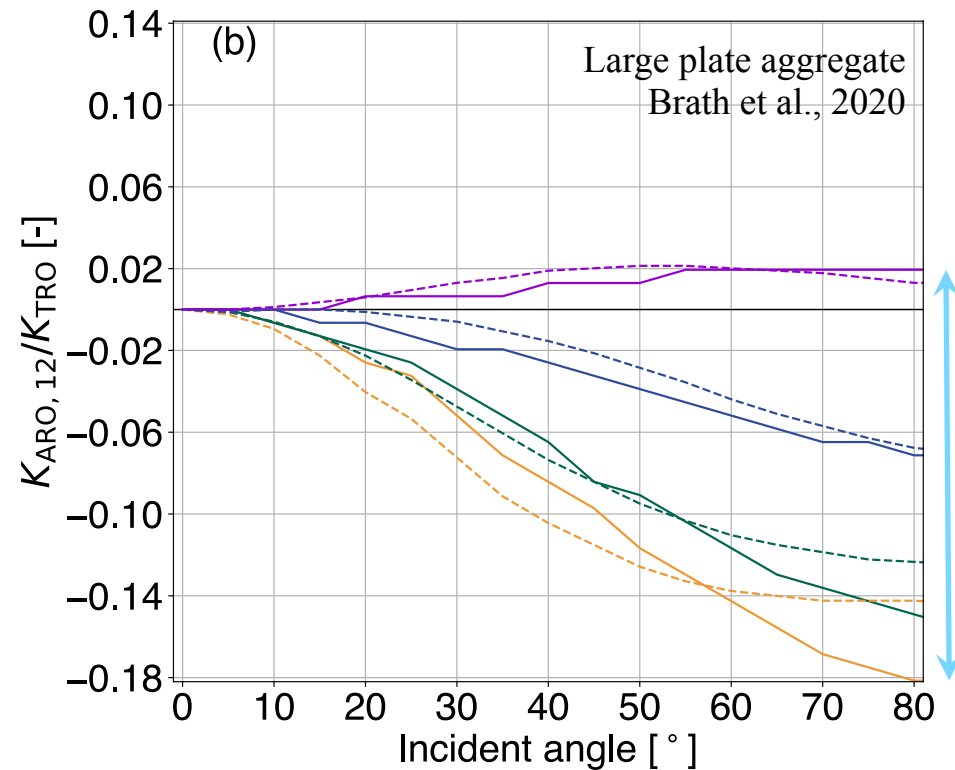
- The parameterization is governed by the polarization ratio ρ :

$$\rho = \frac{\tau_H}{\tau_V} = \frac{\tau_{TRO}(1 + \alpha)}{\tau_{TRO}(1 - \alpha)},$$

α is a correction factor approximating the cross section for linear polarization, i.e., K_{12}

* Control (current framework):

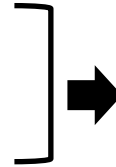
$$\rho (\alpha = 0) = 1$$



● Based on the IFS¹ of ECMWF²

1) Passive monitoring experiments

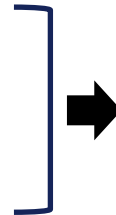
- Sensors: GMI³
- Duration: Jun. – Jul. 2019



Which is the best ρ ?

2) Cycled DA experiments

- Sensors: AMSR-2⁴, GMI³, SSMIS⁵
- Duration: Febr. – May 2019 &
Jun. – Aug. 2019



What is impact of ρ on the forecast?

-
- Microphysical setup following:

Geer, to be submitted to AMT, 2021

¹Integrated Forecast System

²European Centre for Medium-Range Weather Forecasts

³Global Precipitation Mission microwave imager

⁴Advanced Microwave Scanning Radiometer-2

⁵Special Sensor Microwave Imager/Sounder

● Polarization differences (PD)

- Observed (o): $PD^o = T_{BV}^o - T_{BH}^o$
- Simulated (b): $PD^b = T_{BV}^b - T_{BH}^b$

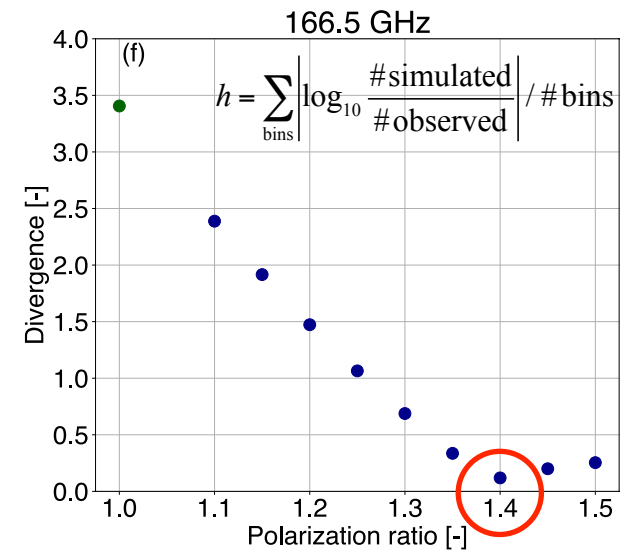
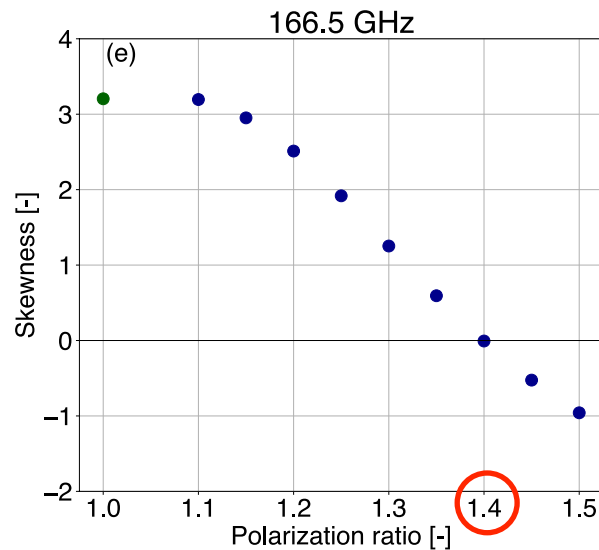
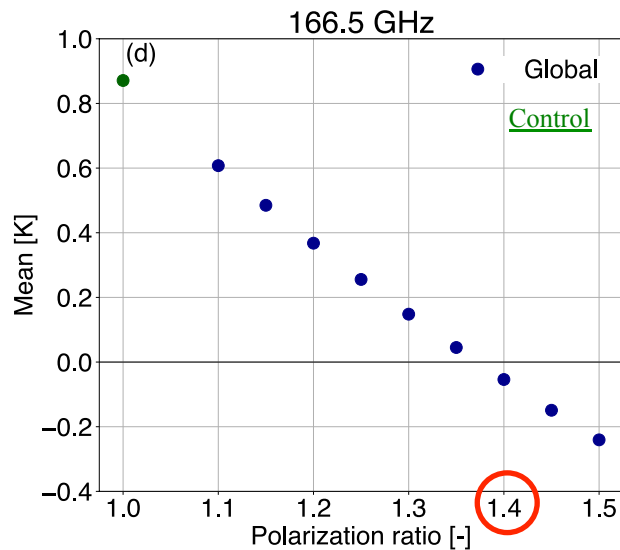
} → $PD^o - PD^b$!

* Careful cloud screening method

● Polarization differences (PD)

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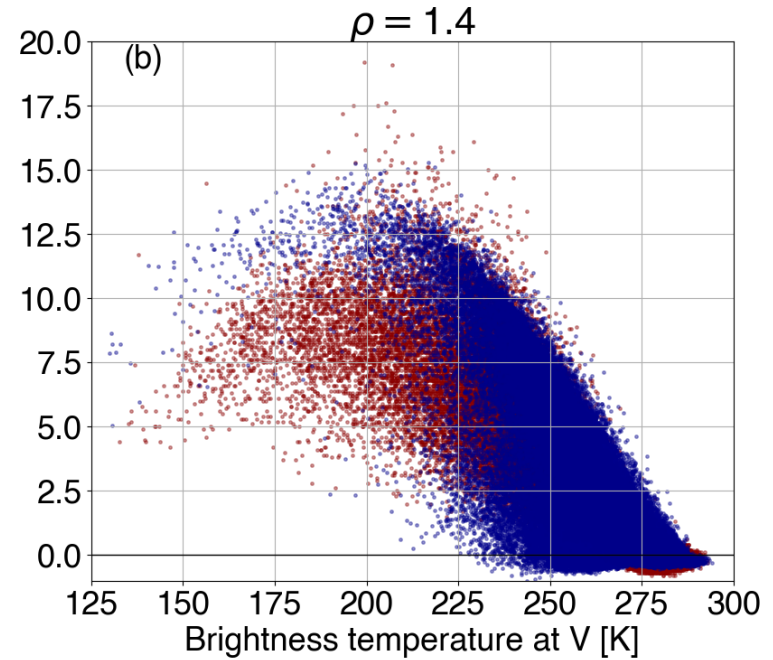
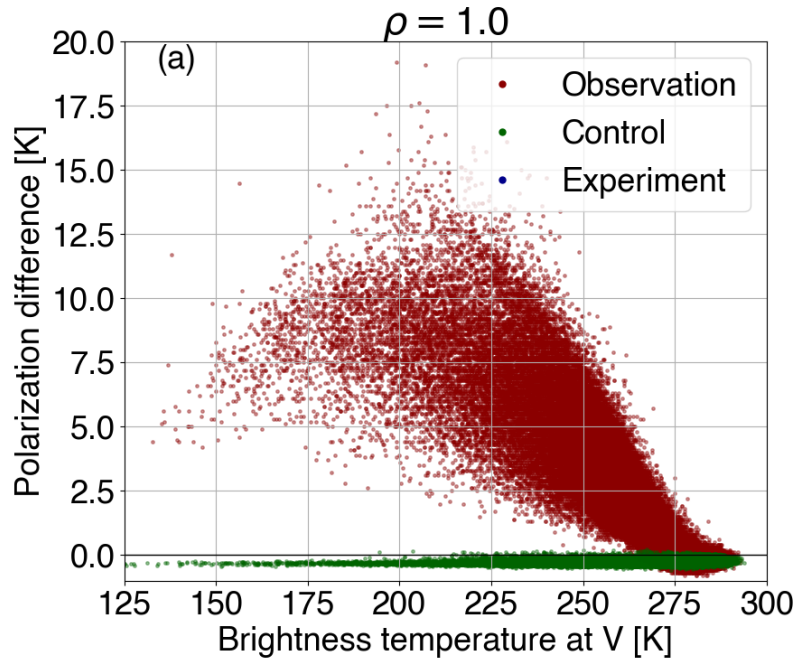
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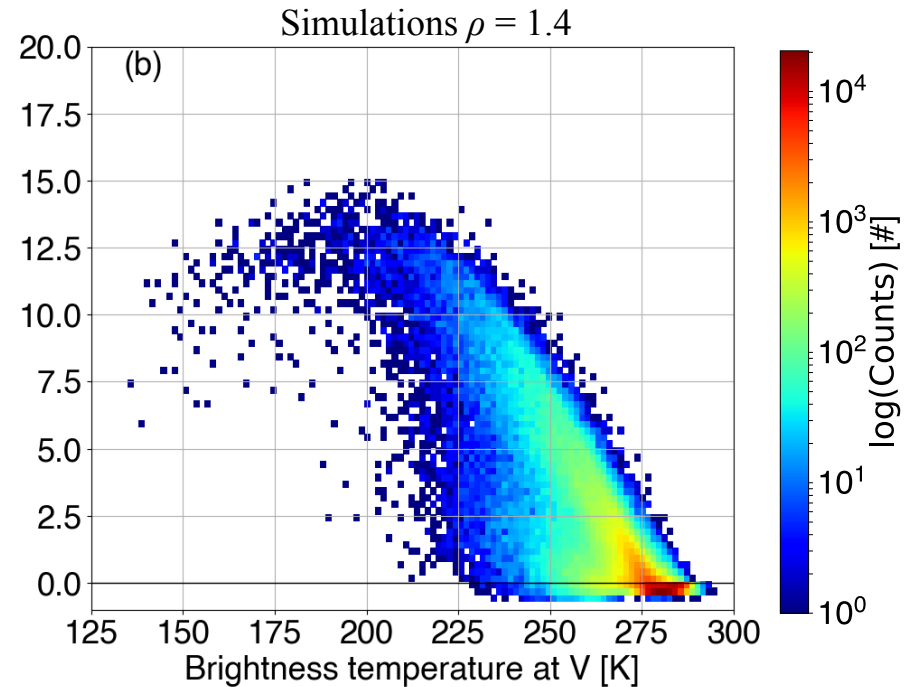
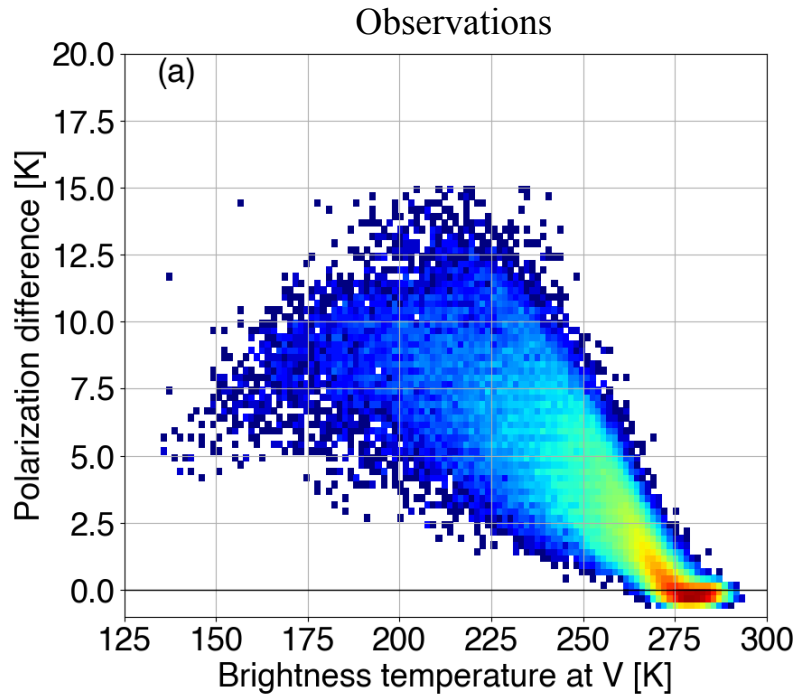
○ Best fit:

166.5 GHz, $\rho = 1.4$ ($\alpha = 0.167$)

89.0 GHz, $\rho = 1.5$ ($\alpha = 0.2$)



- Universal arch-like shape of relation
- Control run completely fails to reproduce it
- A ρ of 1.4 gives a reasonable representation of such relation



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Equal & opposite

$$\alpha = 0.167$$

Unilateral in V

$$\gamma = 0.286$$

$$\rho = \frac{\tau_H}{\tau_V} = \frac{\tau_{\text{TRO}}(1 + \alpha)}{\tau_{\text{TRO}}(1 - \alpha)} = \frac{\tau_{\text{TRO}}(1 + \beta)}{\tau_{\text{TRO}}(1)} = \frac{\tau_{\text{TRO}}(1)}{\tau_{\text{TRO}}(1 - \gamma)} = 1.4$$

Unilateral in H

$$\beta = 0.4$$

¹Advanced Technology Microwave Sounder

²Winds at 850 hPa, AMSUA (53.7H GHz),
geostationary H₂O radiances

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Validation

- Independent: ATMS¹ & conventional² data
 - ☑ Neutral impact

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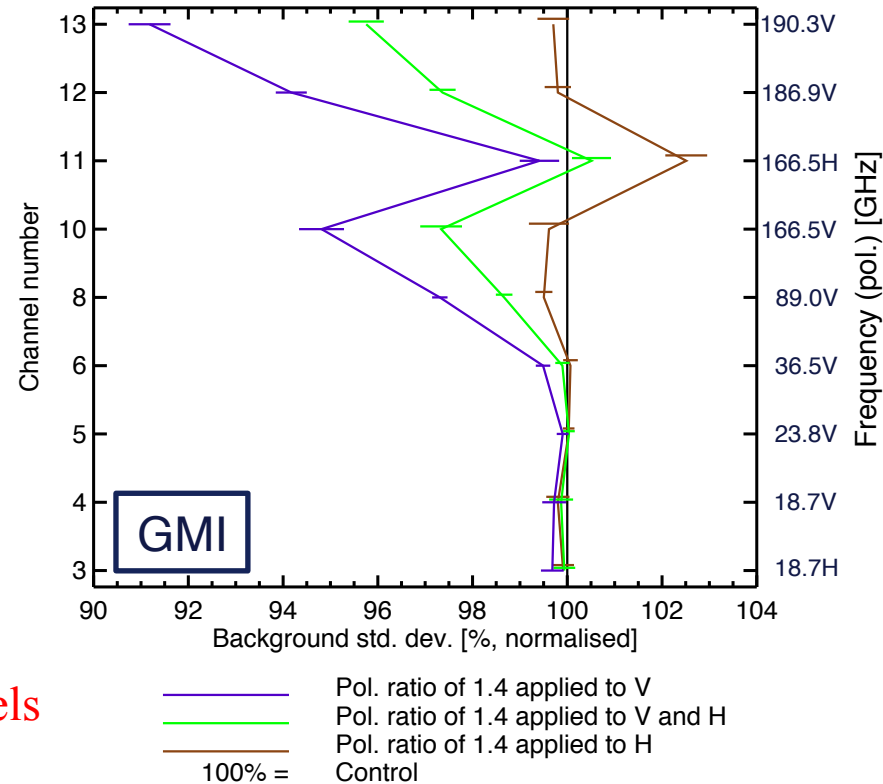
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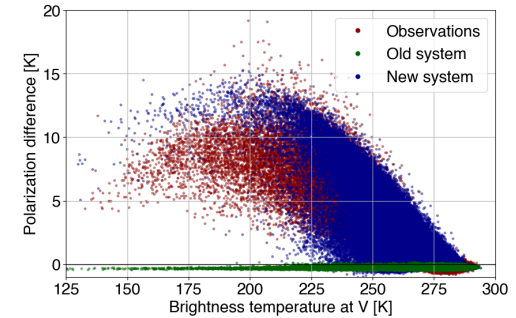
- Independent: ATMS¹ & conventional² data
 - ☑ Neutral impact
 - Dependent: AMSR-2, GMI, SSMIS
 - ☑ Increased consistency between V & H channels
 - ☑ Increased consistency between instruments
- + Tuning the overall level of scattering



¹Advanced Technology Microwave Sounder
²Winds at 850 hPa, AMSUA (53.7H GHz),
 geostationary H₂O radiances

● Final configuration of RTTOV-SCATT v13.0

- Minor software adaption and no calculation burdens
- Maximum modeling errors are reduced by 10–15 K
- Errors are now ~ symmetrical
- Neutral forecast impact, with increasing consistency between instruments
- Polarization ratio of 1.4 in RTTOV-SCATT v13.0 + future IFS cycle



● Outlook

- Polarization correction scheme
 - Cross-track sounders
 - Radar backscattering
 - Ice retrievals, e.g., in GMI
- } 3rd year
- Frequencies above 183 GHz, i.e., Ice Cloud Imager



Photo: esa.int



Photo: Airbus

Model inter-comparison in all-sky millimeter/sub-millimeter radiative transfer

Barlakas et al., *in preparation, JQSRT*, 2021

Objectives:

- Are there any systematic or random errors?
- Where does the two-stream delta-Eddington “falls apart”?
- Most studies focus on clear-sky or frequencies below ~ 90 GHz
- Prepare RTTOV–SCATT for ICI

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ARTS vs RRTOV (–SCATT):

- Clear–sky
- All–sky (i.e. snow, ice, rain)

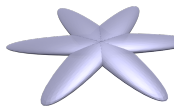
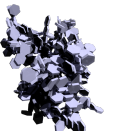
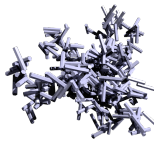
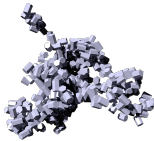
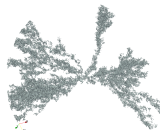
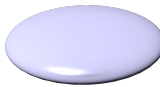
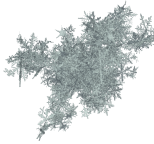
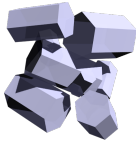
Sensors:

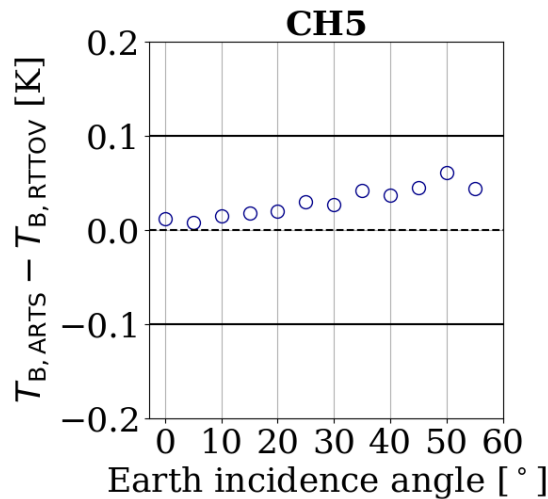
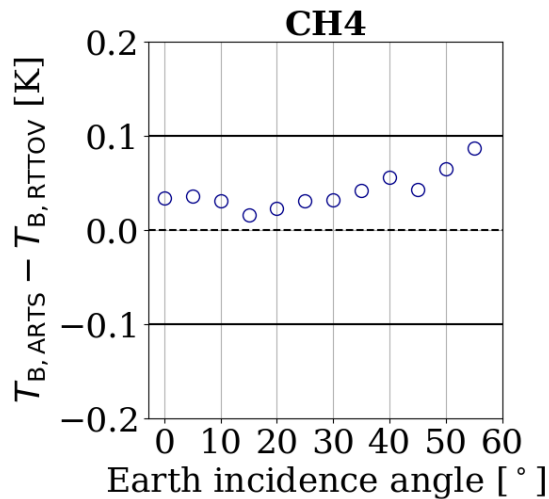
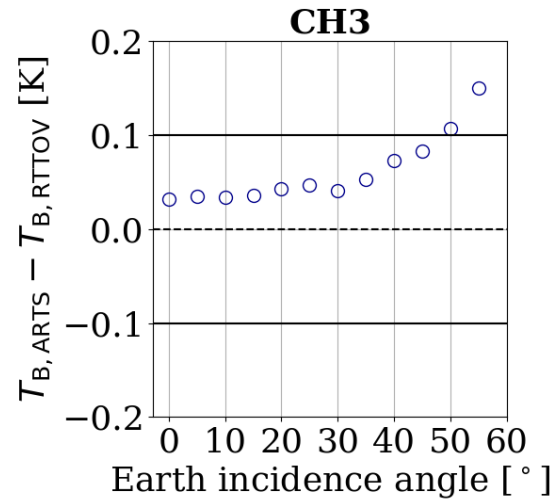
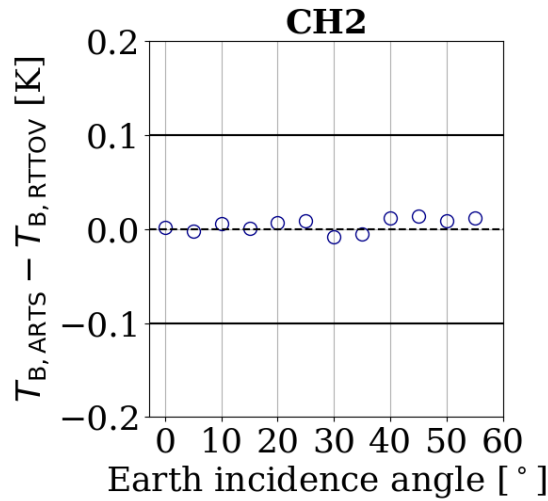
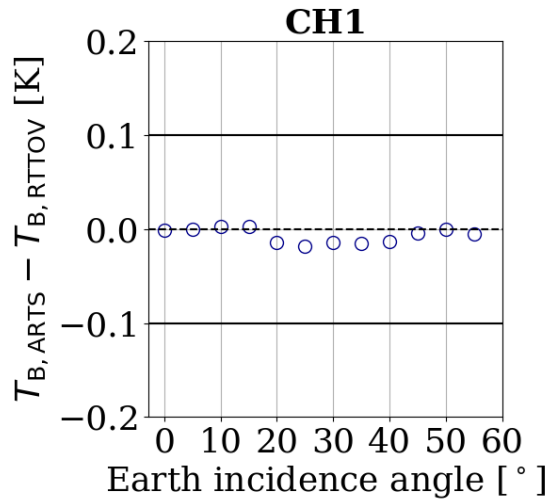
- Advanced Microwave Sounding Unit A (AMSUA)
- Microwave Humidity Sounder (MHS)
- **Ice Cloud Imager (ICI)**

ARTS scattering database

- 34 freq.: 1-886.4 GHz
- 34 particle models (PM)
- 35-45 sizes per PM

Eriksson et al., *ESSD*, 2018





Settings:

- US standard atmosphere
- Surface emissivity = 1

○ An excellent agreement is found ± 0.1 K

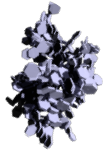
Correct spectroscopy!

○ Exceptions: ICI at 243.2 GHz (~ 0.2 K) and 448 GHz (~ 0.4 K)

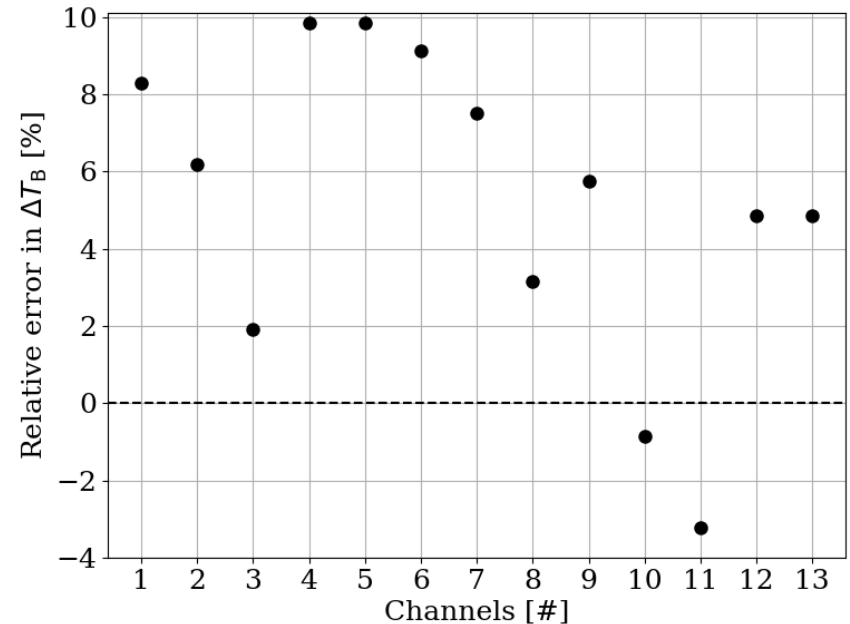
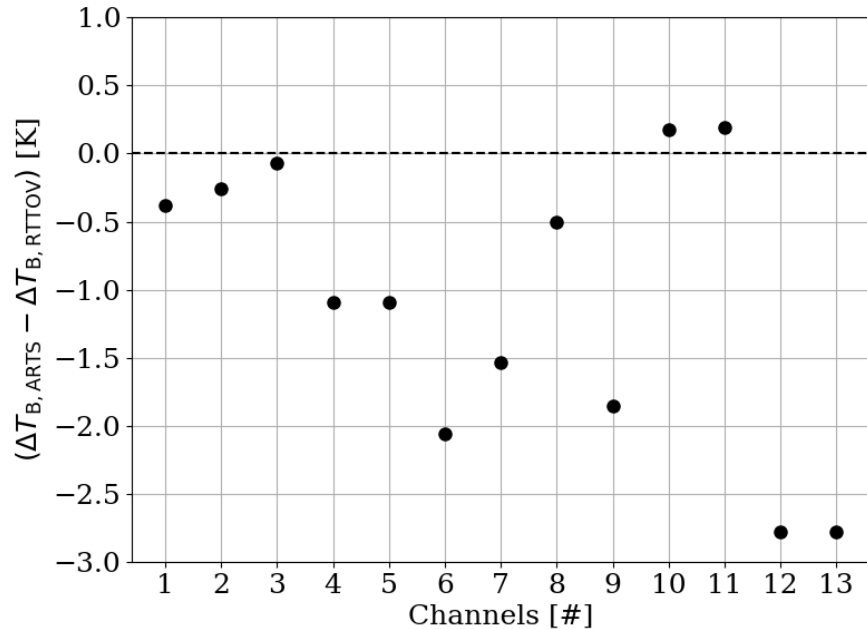
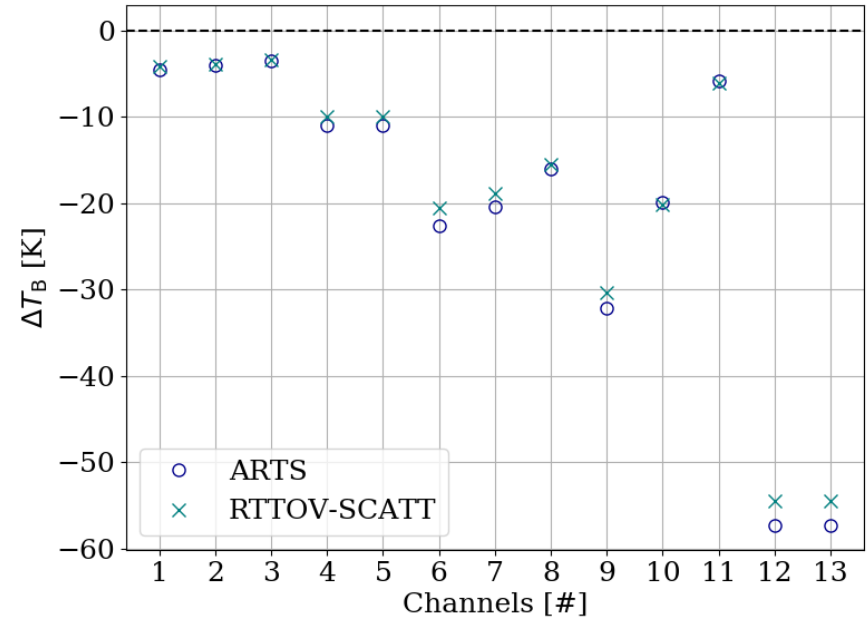
Ongoing by Met Office

Settings:

- Gaussian cloud: 9-11 km
- Large plate aggregate habit
- Field et al., 2007 PSD

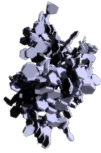


$$\Delta T_B = T_{B,\text{cloudy}} - T_{B,\text{clear}}$$

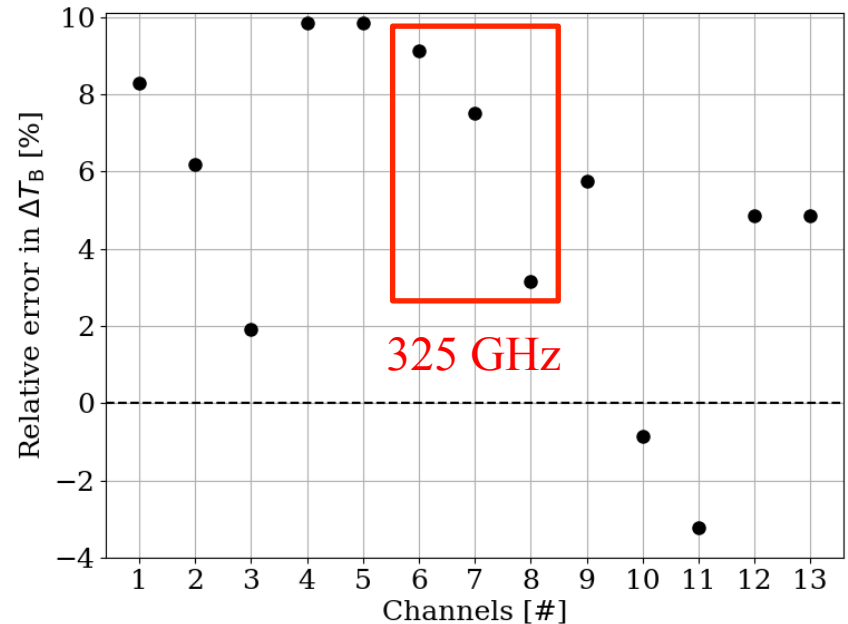
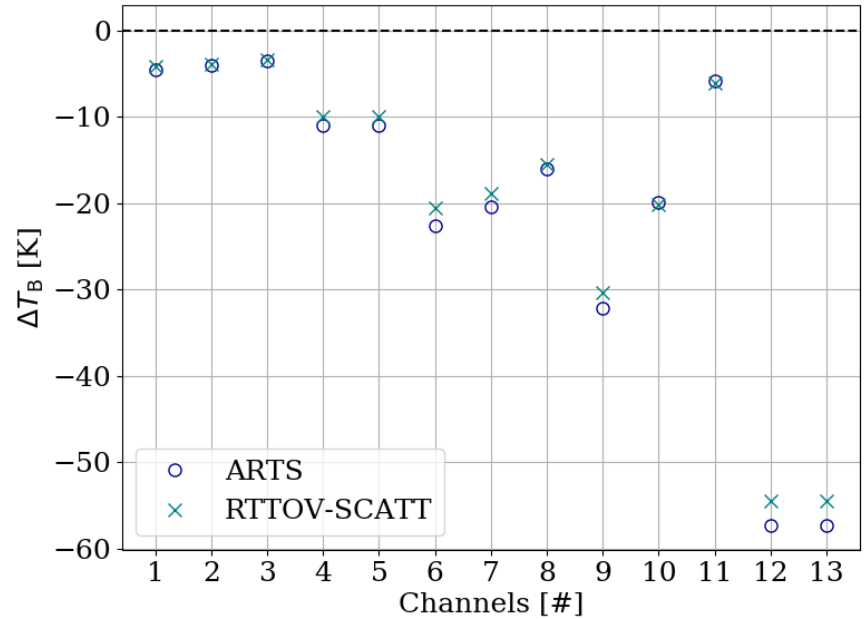
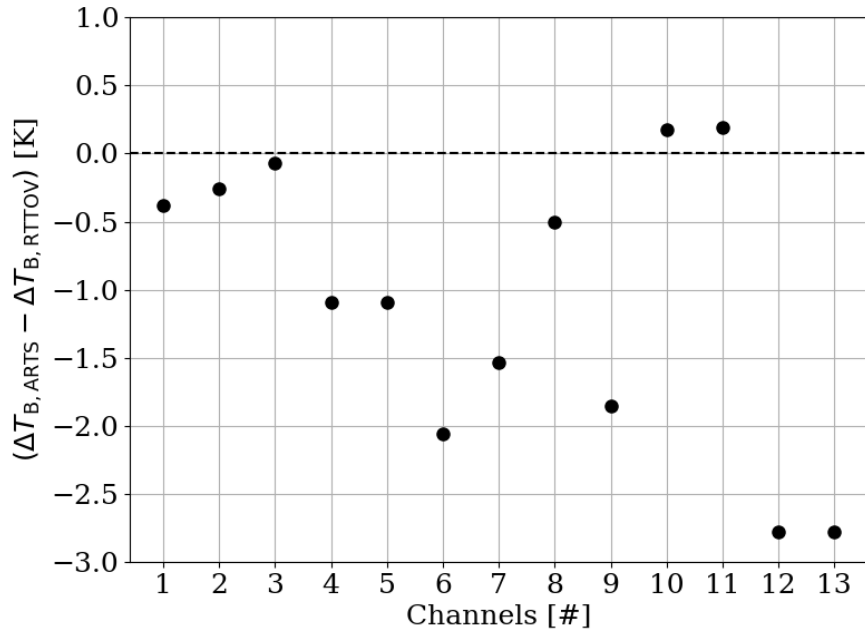


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● Performance of RTTOV-SCATT

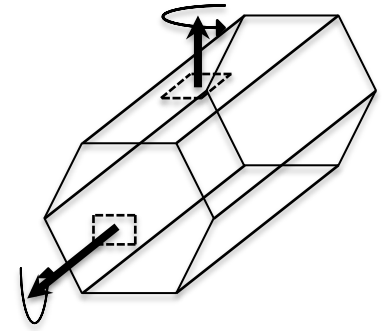
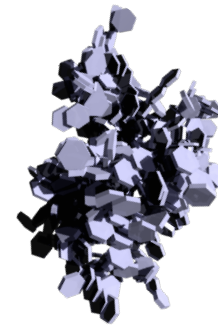
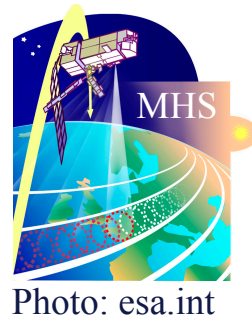
- Clear-sky conditions
 - An excellent agreement is found (± 0.1 K). For ICI, there is still some ongoing work.
- All-sky conditions
 - ICI: $\Delta T_B \sim 3$ K, with a relative error up to $\Delta T_B \sim 10$ %
 - MHS and AMSUA: $\Delta T_B \sim 7$ K, with a relative error up to $\Delta T_B \sim 15$ %

● Outlook

- Where does the two-stream delta-Eddington “falls apart”?
 - Finalize the manuscript
 - Sub-grid variability and cloud overlap scheme.
- } 3rd year

● Polarized correction scheme

- Cross-track sounders
- Radar backscattering
- Ice cloud retrievals (ongoing)



● Performance of RTTOV-SCATT

- Finalize the intercomparison (ongoing)
- Sub-grid variability and cloud overlap scheme.

Thank you for your attention !