## Flat Radial Age Gradients in Massive z~0.6 PSBs



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## The SQuIGGLE Survey We Observe Flat H<sub>s</sub> Profiles – **OIR telescopes: Keck/Gemini/Subaru** Radio/Sub-mm telescopes: ALMA/VLA - 10 $1.0 \cdot$ seconds Η<sub>δ,</sub> The SQuIGGLE Survey is a multi-wavelength study of poststarburst galaxies (PSBs) at z~0.6. See talks by Jenny Greene, -1.0

Justin Spilker, and Wren Suess.

## SQUIGGLE Target Selection and SFH Modeling



-1.5 -1.0 -0.5 0.0 0.5 1.0 -1.5 -1.0 -0.5 0.0 0.5 1.0Arcseconds Arcseconds A subsample our 6 SQuIGGLE galaxies with GMOS IFU follow up in  $H_{\delta}$ . 12 10

seconds

0.5



Annular profiles in the EW  $H_{\delta}$  for

A sample galaxy SQuIGGLE galaxy with rest frame selection filters and the best fit two component star formation history (SFH)

PSBs can be modeled with a composite K+A population.

The K/A ratio and the time since quenching influence the strength of age sensitive features like  $H_{\delta}$ .

our full sample

All our galaxies have EW  $H_{\delta} > 4$  Å at all radii, signifying dominant A-star populations.

## Massive PSBs Quench at All Radii Simultaneously

Setton et al. in prep

Different methods of quenching (e.g. compaction, central starburst) will result in negative, positive, or flat age gradients.

**Ex: Illustris Simulation** 

In contrast to simulations of central starbursts, our galaxies do not show gradients in their radial age.





Wellons et al (2015) radial age profiles from Illustris Simulations If these galaxies are resolved, this indicates that whatever mechanism they quenched by must shut down star formation simultaneously at all radii.

Time since quenching (normalized to the central measurement) fit to our  $H_{\delta}$  measurements by varying K/A and time since quenching