

Review of Polar Stratospheric Cloud Microphysics and Chemistry

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Abstract

The solid and liquid particles which constitute polar stratospheric clouds (PSCs) are of manifold importance to the meteorology of the stratosphere. The heterogeneous reactions which take place on and within these particles release halogens from relatively inert reservoir species into forms which can destroy ozone in the polar spring. In addition, solid PSC particles are instrumental in the physical removal of nitrogen oxides (denitrification) and water (dehydration) of regions of the polar stratosphere. Denitrification, in particular, allows extended ozone destruction by slowing the conversion of chlorine radicals back into reservoir species.

We review the historical development of PSC studies, with particular emphasis on results from the last decade, encompassing developments in observations, in laboratory experiments, and in theoretical treatments. The technical challenge of measuring sufficient of the parameters describing any given polar stratospheric cloud, to allow its microphysics to be understood, has driven forward balloon-borne, aircraft, and satellite instrumentation. The technical challenge of finding suitable laboratory

proxies for PSCs, in order to observe the microphysics under controlled conditions, has resulted in a wide variety of experimental designs, some of which maximise the probability of observing phase change, others of which mimic the surface-volume ratios of PSCs more closely. The challenge to theory presented by PSCs has resulted in improvements in the thermodynamics of concentrated inorganic solutions of volatile compounds, and a new general theory of freezing of water ice from concentrated aqueous solutions. Of the major processes involving PSCs, heterogeneous reaction probabilities for ternary $\text{HNO}_3/\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ solutions, and heterogeneous freezing to produce nitric acid hydrates, are the least well understood.

Key words: Polar stratospheric cloud; stratospheric ozone depletion; stratospheric aerosol; denitrification; heterogeneous chemistry; freezing nucleation

1 Historical overview

2 The first polar stratospheric clouds (PSCs) observed were the beautiful nacre-
3 ous, or mother-of-pearl, clouds which are quite commonly visible above the
4 mountains of Scandinavia and the Antarctic peninsula, but are also occa-
5 sionally seen in the middle latitudes (Stanford and Davis, 1974, simply type
6 “nacreous cloud” into an internet search engine for recent images). For ex-
7 ample, as pointed out to us by an anonymous reviewer, PSCs were recorded
8 in 1901 by the Danish painter Jorgensen. By the mid-20th century, scientists
9 had found that the clouds formed at altitudes of around 25 km (i.e., the lower
10 stratosphere) and temperatures near $-80\text{ }^\circ\text{C}$. Mother-of-pearl clouds are water
11 ice clouds, formed within stratospheric gravity waves. Their rich colouration
12 is the result of forward diffraction of sunlight by the particles; the changes in
13 colour occur because of the variations in particle size along the profile of the
14 wave (e.g., see Tolbert and Toon, 2001, and references within).

1 From satellite and other observations in the last quarter of the twentieth cen-
2 tury, a simple three-stage model of PSC development began to emerge (Figure
3 1). This model starts with the background aerosol particles — consisting of
4 supercooled sulphuric acid solution with about 75% by weight H_2SO_4 (Rosen,
5 1971) — which form the stratospheric Junge layer. As the temperature de-
6 creases these particles may freeze to form sulphuric acid tetrahydrate (SAT)
7 (e.g., see Dye et al., 1992, and references within), although this step was not
8 always included (e.g. Drdla and Turco, 1991). The essential features of this
9 theory were that nucleation of NAT onto the background sulphate particles
10 would occur below the NAT equilibrium temperature ($192 < T_{NAT} < 197 \text{ K}$,
11 for typical partial pressures of water vapour and nitric acid in the lower strato-
12 sphere), leading to Type I PSCs, and further cooling to below the ice frost
13 point ($188 < T_{ICE} < 190 \text{ K}$) would result in the nucleation of water ice onto
14 the NAT particles to form Type II PSCs.

15 However observational data suggested that the formation of PSCs was more
16 complicated. Measurements in the presence of Type I PSCs found that sub-
17 stantial transfer of HNO_3 from the vapour to condensed phase only occurred
18 about 3–4 K below T_{NAT} , and that the HNO_3 partial pressures were signifi-
19 cantly higher than NAT vapour pressures (Rosen et al., 1989; Hofmann et al.,
20 1990; Schlager et al., 1990). These observations suggested that some unidenti-
21 fied process was causing a hysteresis effect, where the formation of NAT was
22 hampered until 3–4 K below T_{NAT} , but, once formed, it could exist up to T_{NAT} .
23 This led to proposals for different requirements for the formation of NAT par-
24 ticles, ranging from the necessity of SAT as a nucleus, to free energy barriers
25 (see Dye et al., 1992, and references within). Rosen et al. (1989) interpreted
26 the difference in vapour pressures as indicating the presence of another meta-

1 stable nitric acid particle, possibly an amorphous solid-solution with variable
2 composition (Hamill et al., 1988). Analysis of lidar measurements (Browell
3 et al., 1990; Toon et al., 1990) identified two distinct subtypes of nitric acid
4 particles. Non-spherical particles, with a spherical volume-equivalent-radius
5 exceeding $1.0 \mu\text{m}$ were denoted as Type Ia, while Type Ib particles were clas-
6 sified as spherical with radii near $0.5 \mu\text{m}$.

7 Hanson (1990) suggested that, when the gas phase has not reached equilib-
8 rium with NAT, supercooled nitric acid solutions could condense up to a few
9 degrees above the frost point. Laboratory experiments on the uptake of HNO_3
10 by $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ solutions led Zhang et al. (1993) and Molina et al. (1993) to
11 conclude that, below T_{NAT} , the background sulphuric droplets would take up
12 H_2O and HNO_3 to form a supercooled ternary solution (STS) particle, com-
13 posed of $\text{HNO}_3/\text{H}_2\text{SO}_4/\text{H}_2\text{O}$, which would freeze to NAT. PSC theory was,
14 therefore, shifting away from the simple three-stage model focused on the solid
15 NAT and ice particles, with other particle phases merely facilitating phase-
16 changes. Thermodynamical modelling studies by Carslaw et al. (1994) and
17 Tabazadeh et al. (1994) demonstrated that STS particles could have HNO_3
18 mass fractions as high as 41%, with H_2SO_4 as low as 3%, in stratospheric con-
19 ditions, provided the particles do not freeze. Tabazadeh et al. (1994) showed
20 that the increase in mass of the particles — due to the uptake of H_2O and
21 HNO_3 — was sufficient to fulfill the size and shape definitions given for Type
22 Ib particles by Toon et al. (1990). Carslaw et al. (1994) compared particle
23 volumes observed by Dye et al. (1992) with those predicted using the differ-
24 ent proposed PSC phases (Figure 2) and demonstrated that PSCs composed
25 of STS particles fitted the data much better than those composed of NAT
26 particles. Bulk freezing experiments (Koop et al., 1995) and theoretical stud-

1 ies (MacKenzie et al., 1995) suggested that homogeneous nucleation rates of
2 NAT from STS were low enough to allow the existence of these liquid particles
3 below T_{NAT} without freezing.

4 The model of PSC development was now a lot more complicated (Figure 3,
5 and Peter (1997)). Three major types of PSC particle had been identified:
6 water ice, NAT and STS. However it was not until five years later that in-situ
7 measurements of particle compositions in the Arctic positively identified the
8 existence of NAT (Voigt et al., 2000a) and STS (Voigt et al., 2000b) in the
9 stratosphere. As the temperature decreases, the background sulphuric acid
10 aerosol particles remain liquid, taking up H_2O and HNO_3 from the gas phase
11 to become STS particles a few degrees below T_{NAT} . However the principal
12 nucleation mechanisms in the stratosphere for converting these liquid particles
13 to solid ice and NAT particles are still not clear.

14 Below, we divide our review of PSC studies into three broad sections: a dis-
15 cussion of the instrumentation deployed in fieldwork, a discussion of PSC
16 composition in detail, and a discussion of the effects of PSCs.

17 **2 Fieldwork Instrumentation**

18 The data used for PSC studies ranges from retrieved atmospheric composition
19 over synoptic scales by satellite-mounted instruments, to the in-situ measure-
20 ment of particle compositions and gas-phase mixing ratios by balloon and
21 aircraft mounted instruments. Most aspects of the stratosphere, from air tem-
22 perature and the mixing ratios of individual components, to the size, number
23 and composition of aerosol particles, can be studied. However, it is not al-

1 ways possible to measure all the atmospheric aspects relevant to a a PSC
2 simultaneously and with the required precision.

3 *2.1 Measuring Aerosol Properties*

4 Even for a homogeneously mixed air parcel, an aerosol population will vary
5 across several different properties, i.e., variations that produce the number,
6 surface area, and volume density distributions, as well as variations in shape,
7 composition and phase (e.g., Jaenicke, 1998; Seinfeld and Pandis, 2006). Each
8 of these characteristics can be represented in a number of different ways, the
9 accuracies of which are usually determined by the limitations of the instru-
10 ments being used. There are few instruments which can measure more than a
11 one or two of these characteristics, and none which will measure all of them
12 (Hinds, 1999; Vincent, 2007).

13 The information content of data from an aerosol instrument is limited by the
14 number of different measurements the instrument makes of the aerosol. For
15 example, a single wavelength lidar will measure the backscatter and depolar-
16 isation caused by an aerosol. From these, and by making some assumptions,
17 we can derive the total aerosol volume concentration ($\mu\text{m}^3 \text{ m}^{-3}$), and an indi-
18 cation of the average particle shape. The total aerosol volume concentration
19 is determined by the number density and size distribution of the aerosol so,
20 by assuming a standard value for one of these characteristics, we can obtain
21 an estimate of the other.

22 To minimise the assumptions required to characterise the aerosol, a number
23 of different instruments, chosen to complement each other, can be used. For

1 instance, it is becoming more common to use data from both optical particle
2 counters and mass spectrometers to obtain data on both the sizes, number, and
3 compositions of the particles (Peter, 1996; Schreiner et al., 1999). However,
4 care must be taken that such combinations are viable. The data collected
5 by instruments provide a snap-shot of the atmospheric conditions. We must
6 make sure that the samples, from which the data are obtained, are related both
7 temporally and spatially, otherwise the information is likely to be misleading
8 when combined.

9 *2.2 Synoptic and Global-Scale Measurements*

10 The first measurements of sub-visible PSCs were made using the SAM-II in-
11 strument, mounted on the Nimbus-7 satellite (McCormick et al., 1981, 1982).
12 SAM-II, and the instruments which followed, are solar occultation photome-
13 ters that measure the attenuation of light from the Sun as it passes through
14 the atmosphere. Lines of sight through the atmosphere to the Sun are only
15 possible during sunrise and sunset, and measurements have sample volumes of
16 several hundred cubic kilometres. Analysis of each measurement is focused on
17 an air parcel centred at the point along the line of sight between the satellite
18 and the Sun which is closest to the surface of the Earth (SAGE III ATBD
19 Team, 2002, available from NASA Langley Research Centre at [http://www-
20 sage3.larc.nasa.gov/](http://www-sage3.larc.nasa.gov/)).

21 The data are processed to obtain the contributions to the solar attenuation
22 of effects such as Rayleigh scattering, aerosol extinction and absorption by
23 gaseous species. To produce vertical profiles, the contributions of the higher
24 altitudes to each measurement are subtracted. SAM-II measured only the $1\ \mu\text{m}$

1 wavelength; however modern instruments take measurements across a number
2 of different wavelengths — the Stratospheric Aerosol and Gas Experiment
3 (SAGE) III instrument produces data over 8 wavelengths between 0.29 and
4 $1.55\ \mu\text{m}$ (SAGE III ATBD Team, 2002), while the Polar Ozone and Aerosol
5 Measurement (POAM) III instrument produces aerosol extinction data for 6
6 wavelengths between 0.353 and $1.018\ \mu\text{m}$ (Strawa et al., 2002).

7 Following McCormick et al. (1981), Poole and Pitts (1994) used the SAM-II
8 aerosol extinction profiles to develop a long term PSC climatology. This in turn
9 has been extended and improved by Fromm et al. (2003) to become a unified
10 stratospheric aerosol and cloud database which combines data from the SAM
11 II, SAGE II, and POAM II/III missions. By taking the ratio of the aerosol
12 extinctions at $0.603\ \mu\text{m}$ and $1.018\ \mu\text{m}$, Strawa et al. (2002) have demonstrated
13 that POAM-III data can be used to distinguish synoptic-scale PSCs of types Ia
14 and Ib (Figure 4). They validated their method using lidar observations that
15 were coincident with POAM observations. Poole et al. (2003), Lee et al. (2003)
16 and Kim et al. (2006) have extended the use of this technique to the analysis
17 of SAGE III measurements and ILAS-I/ILAS-II measurements, respectively.
18 Relative extinction coefficients are calculated using adjacent transmittance
19 profiles, one containing a PSC signature, the second containing none. A non-
20 linear least squares fit method was used to compare the measured extinction
21 spectra with spectra derived using Mie-scattering theory for single (Lee et al.,
22 2003) and dual (Kim et al., 2006) phase PSCs. Using this method they were
23 able to determine, within uncertainties of 25–40%, PSC compositions and ef-
24 fective particle radii.

25 Other limb-sounding instruments, mounted on balloons and satellites, have
26 been also been used to detect PSCs. These instruments scan the atmosphere

1 above the horizon in a similar manner to solar occultation photometers, but
2 are not orientated towards the Sun. Instead they measure the scattering of
3 upwelling tropospheric radiation into the limb view, or the thermal infrared
4 emission from the troposphere, stratosphere and mesosphere. These methods
5 are particularly well suited to PSC studies because they allow measurements
6 to be taken over a greater latitudinal range than the solar occultation method,
7 and also at night (and so within the winter polar regions).

8 Spang et al. (2001) verified a robust cloud detection method of using the ratio
9 of radiances in the wavenumber ranges of 788–796 cm^{-1} and 832–834 cm^{-1} ,
10 which are respectively dominated by CO_2 and aerosol/cloud emissions. Spang
11 et al. (2005) applied this method to the analysis of MIPAS on ENVISAT
12 data, validating the measurements using lidar and satellite data (POAM III
13 and SAGE III). They found that the agreement between the MIPAS and li-
14 dar data was very good. The agreement with the other satellite data was also
15 good, although the solar occultation measurements were able to detect op-
16 tically thinner clouds than the MIPAS measurements. Spang and Remedios
17 (2003) identified a spectral signal at 820 cm^{-1} which they concluded indi-
18 cates the presence of particles with a high HNO_3 concentration (mass fraction
19 $> 40\%$). Höpfner et al. (2006) extended this analysis using refractive indices
20 generated from a number of laboratory studies to generate profiles for the ma-
21 jor PSC particle compositions. Using lidar measurements for validation, they
22 demonstrated that dense homogeneous clouds of ice or small (radii $< 3 \mu\text{m}$)
23 NAT particles can be identified using MIPAS measurements.

24 Satellite-based instruments are important because they allow us to observe
25 large volumes of the atmosphere over several years, so an overview of strato-
26 spheric conditions, and estimations of the extent of PSC coverage, can be

1 made. However, until recently, they could not be used to determine any phys-
2 ical parameters of the aerosols being observed. Even with the information
3 provided by the analysis of extra wavelengths of solar radiation, Strawa et al.
4 (2002) have only been able to distinguish differences in the size of the aerosol
5 particles, and so indirectly determine the type of PSC observed. Because of
6 the scale of these measurements — a typical POAM sample volume is about
7 200 km long, 30 km wide and 1 km thick — these datasets do not capture well
8 small-scale features.

9 *2.3 Mesoscale Instruments*

10 Most remote measurements of PSCs are made using lidar systems, which,
11 unlike passive sensors, contain a laser source as well as an optical sensor. The
12 laser is aimed towards the area of interest, while the sensor detects the light
13 that scatters at 180° , i.e., directly back towards the instrument. The distance
14 between the lidar and the measurement site is calculated from the time the
15 signal takes to return.

16 Two properties of the returned signal are measured, the amount of backscat-
17 tered light and its polarisation, from which are determined the backscatter
18 and aerosol depolarisation ratios. The scattering ratio, S , is the ratio of the
19 total backscatter to the backscatter due to air molecules alone. The ratio of
20 the aerosol backscatter to the molecular backscatter is given by $S - 1$. The
21 total depolarisation ratio, D_T , is the ratio of the perpendicular backscatter
22 return, R_\perp , to the parallel backscatter return, R_\parallel

$$23 \quad D_T = \frac{R_\perp}{R_\parallel} . \tag{1}$$

1 The aerosol depolarisation ratio is (Browell et al., 1990)

$$2 \quad D_a = D_T \frac{S_{\parallel}(S_{\perp} - 1)}{S_{\perp}(S_{\parallel} - 1)}, \quad (2)$$

3 where S_{\perp} and S_{\parallel} are the the scattering ratios for the perpendicular and parallel
4 returns, respectively.

5 Lidar measurements can be made using a single laser wavelength (e.g., see
6 Poole and McCormick, 1988; Gobbi et al., 1998; Santacesaria et al., 2001),
7 however the amount of information that can be gleaned from the data is
8 greatly increased if two or more wavelengths of light are used (Browell et al.,
9 1990). This allows the use of the colour ratio

$$10 \quad C(\lambda_2, \lambda_1) = \frac{S(\lambda_2) - 1}{S(\lambda_1) - 1}. \quad (3)$$

11 Two typical wavelengths used for lidar studies are $\lambda_2 = 1.064 \mu\text{m}$ and $\lambda_1 =$
12 $0.603 \mu\text{m}$. The wavelength dependence of the aerosol backscatter is assumed to
13 be proportional to $\lambda^{-\alpha}$, while molecular scattering is proportional to λ^{-4} . Thus
14 the wavelength dependence of the aerosol backscatter ratio is proportional to
15 $\lambda^{4-\alpha}$. The aerosol backscatter wavelength dependence parameter α is given by

$$16 \quad \alpha = 4 - \frac{\ln(C(\lambda_2, \lambda_1))}{\ln(\lambda_2/\lambda_1)}, \quad (4)$$

17 and is size dependent. A large α indicates that a majority of particles have
18 radii smaller than the lidar wavelengths, while a small α indicates that the
19 majority of particles have radii larger than the lidar wavelengths (Toon et al.,
20 2000; Shibata et al., 1997).

21 The aerosol backscatter depends on the particle size distribution, particularly
22 the surface area density, and the particle refractive index, as well as the lidar

1 wavelength. The amount of depolarisation occurring depends on the shape of
2 the particles. It is small for molecular scattering, and zero for spherical parti-
3 cles, but will occur with even slight departures from spherical shapes. While
4 Mie scattering theory can be used to analyse PSCs consisting of spherical
5 particles (e.g., see Toon et al., 2000), detailed analysis of lidar observations of
6 many PSCs is difficult because they consist of nonspherical solid particles with
7 sizes comparable to the wavelength of the laser. Theoretical modelling can aid
8 analysis; however this must be done by solving the Maxwell equations using
9 numerical techniques (see Liu and Mishchenko, 2001, and references within).

10 Analysis of lidar data requires knowledge of the purely-molecular backscatter
11 (Browell et al., 1990); lidar measurements are often normalised, therefore, at a
12 high altitude where the signal is assumed to be from molecules alone. However,
13 as discussed by Toon et al. (2000), stratospheric aerosols are ubiquitous, and
14 so data used for normalisation may not be purely molecular. They also noted
15 that aerosol scattering is often not much larger than molecular scattering,
16 and so only considered data for which $S(0.603) > 1.2$ and $S(1.064) > 2.0$ to
17 contain PSCs.

18 Using the method described above, Browell et al. (1990) identified the exist-
19 ence of two different types of aerosols above the water ice frost point, Types
20 Ia and Ib, and produced a lidar-based classification (Table 1). Toon et al.
21 (1990) used these observations to describe several physical parameters of Type
22 I clouds. The value of α puts limits on the particle size: large values (Type Ib)
23 indicated particles with mode radii less than the laser wavelength, i.e. about
24 $0.5 \mu\text{m}$. Smaller α values (Type Ia) limit the mode radii of the observed par-
25 ticles to $1 \mu\text{m}$ or greater. The amount of depolarisation indicates the shape
26 of the particles: the very low depolarisation of the Type Ib particles suggests

1 that they are almost spherical, while the higher values for the Type Ia particles
2 indicates that they are much less spherical. The backscatter ratio can be used
3 to calculate the total volume concentration of particles. The high value for the
4 Type Ib particles indicates that they have a large volume concentration, and
5 are likely to have taken up most of the atmospheric HNO_3 . Conversely, Type Ia
6 particles are likely to have a small total volume concentration. However some
7 non-spherical particles scatter much less efficiently than spheres at 180° so this
8 relationship is less clear cut. Although work has been done on characterising
9 the lidar response to non-spherical particles (Liu and Mishchenko, 2001) and
10 aerosols composed of a number of different particle shapes (Reichardt et al.,
11 2002), this is still a complicated and uncertain field.

12 Lidar measurements do not provide the compositions of the particles that they
13 detect. To deduce PSC composition, von König et al. (2002) combined lidar
14 measurements of PSCs with gas-phase HNO_3 measurements from the Air-
15 bourne Submillimeter Radiometer (ASUR). The ASUR instrument measures
16 radiation at wavelengths of about 0.5 mm, large enough to not be affected by
17 scatter from stratospheric particles, so in the stratosphere the signal is only
18 influenced by absorption by gaseous molecules. By combining the gas-phase
19 HNO_3 measurements with temperature and H_2O data, and lidar observations
20 of PSCs, von König et al. (2002) were able to conclude that the observations
21 in the Arctic on one particular day were consistent with a NAT composition.
22 These observations are limited by the ASUR instrument, which has a vertical
23 resolution of 6–10 km, a horizontal resolution of about 20 km, and a precision
24 of about 0.3 ppb in the lower stratosphere (see von König et al., 2002, and
25 references within). This resolution is good enough for synoptic-scale PSCs,
26 but would not be able to distinguish the thin layering (50–300 m thick) which

1 has been observed within mesoscale PSCs (Hofmann, 1990).

2 Remote-sensing instruments can provide measurements of particle shapes and
3 sizes as well as of the total aerosol volume concentration. Combining these
4 with measurements of gas-phase constituents and standard stratospheric num-
5 ber densities allows particle compositions and size distributions to be inferred.
6 However, these measurements are limited by low spatial resolution. For testing
7 microphysical models, the lack of direct composition information is a disad-
8 vantage of lidar measurements, but the detailed spatial information is a great
9 advantage (e.g., Tsias et al., 1999; Wirth et al., 1999).

10 *2.4 In-situ Instruments*

11 In-situ instruments cannot provide the same large-scale coverage as remote
12 sensing instruments. However, they do often provide a higher spatial resolu-
13 tion, ca. 100-500m horizontally, allowing the study of small-scale features. To
14 get these instruments up to the stratosphere they can be mounted either on
15 balloons (e.g., Hofmann and Deshler, 1990; Schreiner et al., 2003) or high-
16 altitude aircraft, such as the American ER-2 (e.g., Dye et al., 1992; Northway
17 et al., 2002) and the Russian M-55 Geophysica (e.g., Stefanutti et al., 1999).
18 These platforms have different advantages and disadvantages. Aircraft are very
19 flexible, allowing focused studies of small-scale features. However they have a
20 limited range. Balloons have a potentially far greater range than aircraft, and
21 follow near-lagrangian flight paths, enabling the study of the evolution of air
22 parcels in time. However they can carry fewer instruments, have limited op-
23 tions for steering, and can only be launched when ground-level weather is near
24 calm.

1 The in-situ instruments most similar to the remote-sensing instruments de-
2 scribed above are the backscatter sondes. Backscatter sondes are like lidar
3 but without range detection (i.e. all the backscattered photons from a flash
4 are counted together), and are used to study particles within about 50 metres
5 of the platform. These use backscatter, depolarisation and colour ratios to
6 measure particle shapes and total aerosol volume concentrations in the same
7 way as the lidar systems described above (e.g., see Schreiner et al., 2003; Lowe
8 et al., 2006, and references within).

9 Optical particle counters, such as the Multiangle Spectrometer Probe (MASP)
10 and Forward Scattering Spectrometer Probe (FSSP), are used to measure the
11 number density and size distribution of the aerosol population. These instru-
12 ments measure the scattered light produced when individual particles pass
13 through a beam of light (which can be either a laser or white light). The
14 inlets of instruments mounted on aircraft are orientated so that the atmo-
15 sphere can be passively sampled using the movement of the aircraft, allow-
16 ing the particles to be measured under near-ambient conditions. Instruments
17 mounted on balloons cannot do this, and so must use pumps to collect sam-
18 ples. Light is scattered in all directions by the particles. All instruments collect
19 the strong forward-scattered signal, though some instruments also collect the
20 weaker back-scattered signal (e.g. MASP, see Drdla et al., 2003, and references
21 within). Mie scattering theory is used to calculate particle sizes, with the as-
22 sumption that the sampled particles are spherical. In addition, a standard
23 refractive index must be used, usually set to a value between 1.40–1.45, based
24 on an estimation of the average refractive index of stratospheric aerosols. These
25 instruments can typically resolve particles between $\sim 0.2 \mu\text{m}$ and $\sim 20 \mu\text{m}$ in
26 size. It is also possible to obtain particle number densities. The main limiting

1 factors on accuracy are uncertainties associated with the sample volume. A
2 statistically significant sample needs to be about 100 particles or more, so
3 sample volumes need to be large to measure the diffuse aerosols present in
4 the stratosphere. This means that sample times can be up to 2 minutes,
5 resulting in a spatial resolution of 10–20 km for aircraft mounted instruments
6 (Baumgardner et al., 1992).

7 A chemiluminescence detector for gas-phase NO_y and NO has been mounted
8 on the ER-2 for polar missions since the late 1980s (Fahey et al., 1989). NO_y
9 is the sum of all nitrogen oxides excluding N_2O , i.e., NO, NO_2 , HNO_3 , etc. In
10 the wintertime lower stratosphere, the majority of NO_y is HNO_3 . Recently the
11 instrument has been adapted for measuring large HNO_3 -containing particles
12 (Northway et al., 2002). This instrument, which has subsequently been dupli-
13 cated on the Geophysica (Voigt et al., 2005), has two inlets. The front inlet
14 measures the NO_y and NO content of the gas-phase and all sizes of particles.
15 The rear inlet makes use of particle inertia to avoid sampling any particles
16 larger than $2\ \mu\text{m}$ aerodynamic diameter. Subisokinetic inlets are used — the
17 internal flow rates are constant and lower than the true speed of the air-
18 craft — which enhance particle number concentrations within the instrument.
19 The factor of enhancement can be calculated using fluid dynamics simula-
20 tions (Northway et al., 2002), and increases with increasing particle size. By
21 subtracting the NO_y measurement of the rear inlet from that of the front in-
22 let the NO_y content of the large ($> 2\ \mu\text{m}$) particles can be measured. When
23 these particles are collected at a lower frequency than the sampling rate (1 Hz)
24 then individual particles can be identified. The size of the individual particles
25 can be calculated using their net NO_y content and assuming a set chemical
26 composition (usually NAT).

1 The compositions of smaller particles can be measured using the balloon
2 mounted Aerosol Composition Mass Spectrometer (ACMS) (Schreiner et al.,
3 2002). In stratospheric conditions this instrument has a throughput of around
4 $15 \text{ cm}^3 \text{ s}^{-1}$. The sampled air passes through an aerodynamic lens — which
5 concentrates particles within the size range $0.3\text{--}5 \mu\text{m}$ into a narrow beam —
6 into two vacuum chambers which reduce the air pressure in two steps down to
7 10^{-8} mbar, used to produce an enhanced aerosol-to-gas signal ratio within the
8 mass spectrometer. Particles take about 0.5 ms to pass through the vacuum
9 chambers. Model calculations and experimental results show that the change
10 in composition of particles larger than $0.1 \mu\text{m}$ during this time is negligible
11 (Schreiner et al., 2002). From the vacuum chambers the aerosol beam passes
12 into an evaporation sphere, which is heated to 80°C . Finally the majority of
13 the resulting gas enters the mass spectrometer to be measured, while approx-
14 imately 30% of the particle mass is lost back to the vacuum chamber. The
15 background signal of the instrument due to the remaining part of the gas-phase
16 can be determined by moving the focus of the lens away from the evaporation
17 sphere, stopping any particles from entering the sphere and so being measured.
18 The relative count rates of the different chemical species measured, typically
19 H_2O and HNO_3 , allowing particle composition to be determined. However the
20 sizes and number of particles measured cannot be reliably determined due to
21 the mixing of the signals from different particles and the loss of mass from the
22 evaporating sphere.

1 **3 Particle Compositions**

2 Since the discovery of PSCs, a number of different types of particles have
3 be postulated as as being important in the formation and lifecycles of these
4 aerosols. Observations of PSCs have been split into Types Ia, Ib and II based
5 on classifications of lidar measurements (Poole and McCormick, 1988; Toon
6 et al., 1990). However these divisions are based on the optical parameters
7 of the aerosols, which, while providing an idea of typical shapes and sizes
8 of particles, contain very little information about the actual composition of
9 the particles. Below we will discuss the current ideas on the compositions of,
10 and transitions between, the different types of PSC particles. The broad lidar
11 classifications will be used to provide a loose structure; however on occasions
12 we will stray outside these limits.

13 *3.1 Type Ib particles*

14 Because of their particle sizes and shapes, Type Ib PSCs are believed to consist
15 mostly of STS droplets. These are formed by the condensational growth of the
16 background stratospheric aerosol, which is comprised of binary $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$
17 particles (Rosen, 1971). The amount of sulphuric acid in the condensed phase
18 increases significantly as the temperature decreases from ~ 240 to ~ 210 K
19 (Figure 5). For temperatures of 200 K and below it can be assumed that all
20 the sulphuric acid is in the condensed phase. This makes changes in the com-
21 positions of the binary particles easier to calculate, because only transport
22 of water from/to the vapour phase need be considered. The mass fraction of
23 H_2SO_4 in the droplets decreases with decreasing temperature, because of the

1 extra water vapour which is taken up by the particles. The first parameteri-
2 sation for the composition and volume of stratospheric $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ aerosols
3 was developed by Steele and Hamill (1981).

4 There are several other gaseous species which will dissolve into the strato-
5 spheric $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ droplets, including the halogen “reservoir” species: HCl,
6 HBr, HOCl and HOBr. Halogen compounds may react in the liquid aerosol
7 to release photochemically active species which catalytically destroy ozone
8 (Solomon, 1990, 1999). Thermodynamical models have been developed to
9 model the equilibrium properties of liquid stratospheric droplets (e.g. Carslaw
10 et al., 1994; Cox et al., 1994; Tabazadeh et al., 1994). The models show a
11 rapid change in particle composition about 4–5 K above the ice frost point,
12 from almost binary $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ solutions to almost binary $\text{HNO}_3/\text{H}_2\text{O}$ so-
13 lutions (Figure 6). At these temperatures, the majority of HCl and HBr in
14 the system transfers to the condensed-phase (Fig. 6a). However these species
15 are much less abundant than HNO_3 , and so are only minor constituents of
16 droplets (Fig. 6b).

17 Particles do not always maintain equilibrium with the atmosphere. For ex-
18 ample, Peter et al. (1994) noted that NAT particles take several hours to
19 re-equilibrate with the gas-phase after undergoing a rapid change in condi-
20 tions. In the case of crystalline particles, such as NAT or SAT, with a fixed
21 composition, this process merely entails a change in the mass of the particle.
22 Liquid particles, on the other hand, do not have a fixed composition. Thus, in
23 theory, they can adopt very different compositions during this re-equilibrium
24 process from that which would be the case if equilibrium was maintained. Us-
25 ing a microphysical model, Meilinger et al. (1995) investigated the behaviour
26 of an ensemble of STS droplets subjected to rapid temperature changes. They

1 found that, during rapid cooling, the majority of the particles in the distribu-
2 tion would not take up HNO_3 rapidly enough to remain in equilibrium with
3 the atmosphere. The resulting increase in the HNO_3 saturation ratio allows
4 the smallest particles to overcome the Kelvin effect and rapidly grow, gaining
5 a far higher HNO_3 mass fraction than would otherwise occur (Figure 7). The
6 hysteresis effect observed in the evolution of the HNO_3 mass fraction occurs
7 because HNO_3 is a factor of ~ 500 less abundant than H_2O in the stratosphere.
8 The time taken for a particle constituent to regain equilibrium with the gas-
9 phase is dependent on the particle size and partial pressure of the constituent.
10 Thus, after an infinitesimal temperature change, the time taken for HNO_3 in
11 a particle to regain equilibrium with the gas-phase is 7 hr and 10 min, for par-
12 ticles of radius 2 and $0.1 \mu\text{m}$, respectively. For H_2O the time lag decreases to
13 29 s and 0.7 s, respectively (Meilinger et al., 1995).

14 Figure 8 demonstrates the importance of non-equilibrium compositions. The
15 data shown (diamonds) are from a balloon flight from Kiruna, Sweden, on 25
16 January 1998 (Schreiner et al., 1999), and show $\text{HNO}_3 : \text{H}_2\text{O}$ mole ratios from
17 mass-spectrometer measurements in a PSC. The data clearly do not fall on the
18 equilibrium STS composition, even considering a range of possible gas-phase
19 abundances of HNO_3 . Using a non-equilibrium model for the growth of the
20 liquid PSC droplets (Lowe et al., 2003) captures much more of the observed
21 variability on composition. As expected from the relaxation time arguments,
22 above, the non-equilibrium particles are usually deficient in HNO_3 , by up
23 to a factor of three at $T - T_{ice} \approx 4\text{K}$, for example. Such disequilibrium in
24 mountain-wave PSCs has also been observed in aircraft measurements (Lowe
25 et al., 2006).

26 SAT particles are not directly related to observations of Type Ib PSCs; they

1 are, however, the thermodynamically stable form of $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ particles be-
2 tween 188–213 K (Koop and Carslaw, 1996). Thus we must consider their pos-
3 sible effects on the growth of STS droplets from the background sulphate par-
4 ticles. SAT is a crystalline solid with the chemical composition $\text{H}_2\text{SO}_4 \cdot 4\text{H}_2\text{O}$.
5 It is part of a broad family of crystalline sulphuric acid hydrates, which
6 includes the monohydrate, SAM ($\text{H}_2\text{SO}_4 \cdot \text{H}_2\text{O}$), the hemihexahydrate, SAH
7 ($\text{H}_2\text{SO}_4 \cdot 6.5\text{H}_2\text{O}$) and and the octahydrate, SAO ($\text{H}_2\text{SO}_4 \cdot 8\text{H}_2\text{O}$).

8 Koop and Carslaw (1996) studied the stability conditions for H_2SO_4 solid
9 phases by comparing the composition of stratospheric $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ droplets
10 with the melting points of the different solid phases (Figure 9). A useful mea-
11 sure of the stability of the solid phases is their saturation ratios in an atmo-
12 sphere of the relevant composition. When these are greater than 1 then the
13 solid is stable, less than 1 and the solid is unstable. From Figure 9B we would
14 deduce that SAT particles form ice particles below 188 K (point c). Figure 9A
15 also suggests that SAT particles should change to different solid phases when
16 the temperature increases above point (a). However, laboratory experiments
17 have shown that SAT will melt rather than change to a different solid phase
18 when warmed above 210–215 K (Middlebrook et al., 1993).

19 As discussed above, liquid stratospheric aerosols absorb large amounts of
20 HNO_3 at low temperatures (Zhang et al., 1993; Molina et al., 1993), which will
21 change the behaviour of stratospheric SAT particles when cooled. The stability
22 of dry SAT particles can be determined by examining the saturation ratio in
23 liquid droplets with the same composition as SAT under the same atmospheric
24 conditions (Figure 10). The addition of HNO_3 to the system significantly re-
25 duces the stability of SAT, which deliquesces simply because it is less thermo-
26 dynamically stable than STS. However experimental results (Iraci et al., 1998)

1 indicate that this process may be kinetically limited. Only $\sim 60\%$ of SAT sam-
2 ples subjected to conditions below the critical temperature of deliquescence
3 (point 1 in Figure 10) exhibited SAT loss to form $\text{HNO}_3/\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ solution.
4 The composition and stability conditions for SAT, as well as its phase tran-
5 sitions to aqueous H_2SO_4 and HNO_3 solutions, are well known. However the
6 formation routes for SAT are more obscure. Laboratory experiments (Middle-
7 brook et al., 1993; Beyer et al., 1994; Koop et al., 1997b) have not observed the
8 homogeneous nucleation of SAT above the ice point, which, with the addition
9 of an analysis of nucleation statistics (Koop et al., 1997b), suggests that sul-
10 phate aerosols do not freeze homogeneously to form sulphuric acid hydrates.
11 Laboratory experiments (Bogdan et al., 2003) also found that concentrated
12 (>40 wt% H_2SO_4) sulphate droplets will not heterogeneously freeze on sil-
13 ica particles at stratospheric temperatures. It has been suggested that liquid
14 $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ droplets may behave like glassy solids below ~ 194 K and crys-
15 tallize upon warming to form SAT (Tabazadeh et al., 1995). However Koop
16 et al. (1997b) showed that the rate of crystal growth is such that stratospheric
17 aerosol droplets would freeze as soon as nucleation has occurred. In labora-
18 tory work, SAT is formed using supercooling, a process which invariably also
19 produces water ice. Thus it is likely that SAT heterogeneously nucleates on
20 other crystalline solids in the stratosphere; however no studies of nucleation
21 rates have been made.

22 3.2 *Type II particles*

23 Type II PSCs were identified as consisting of large ($\geq 10 \mu\text{m}$) crystalline parti-
24 cles which only exist below the stratospheric ice frost point of ~ 188 K (Poole

1 and McCormick, 1988). The obvious choice of composition for these particles
2 is water ice (Turco et al., 1989). However, because these ice crystals nucleate
3 from HNO₃-rich particles, the compositions and lifecycles of these particles
4 must be more complicated than those of pure H₂O ice crystals.

5 Three routes of ice formation have been suggested: vapour deposition of ice
6 onto either SAT or NAT, homogeneous nucleation from an aqueous solu-
7 tion, and heterogeneous nucleation in aqueous solutions containing solid cores.
8 Vapour deposition of ice onto NAT particles was postulated as a method of
9 formation for Type II particles while the three-stage model of PSC forma-
10 tion was in favour (e.g. Hanson and Mauersberger, 1988; Peter et al., 1991).
11 However no studies have been made into this route of formation. Recent ex-
12 perimental work into the vapour deposition of ice on to SAT has observed ice
13 formation at supercoolings of ≤ 1 K below the ice frost point (Fortin et al.,
14 2003). However this route would require the existence of SAT below its deli-
15 quescence temperature. This may occur for up to $\sim 40\%$ of the available SAT
16 particles (Iraci et al., 1998), but is more likely when the majority of the avail-
17 able HNO₃ has been taken up into NAT particles (lowering the deliquescence
18 temperature).

19 With the addition of STS to the list of possible PSC particles, the homo-
20 geneous nucleation of ice from STS was proposed. Laboratory experiments
21 with solutions such as H₂SO₄/H₂O, HNO₃/H₂O and HNO₃/H₂SO₄/H₂O have
22 shown that homogeneous nucleation of ice requires the solution to be super-
23 cooled 2–3 K below the ice frost point (e.g., Chang et al., 1999; Koop et al.,
24 1998; Middlebrook et al., 1993). Theoretical studies, using classical homoge-
25 neous freezing theory, are consistent with this (MacKenzie et al., 1995, 1997,
26 1998). Using experimental data Koop et al. (2000) developed a thermody-

1 namic scheme for homogeneous ice nucleation, expressed as a function of the
2 water activity and pressure. These experiments and theories have assumed
3 that the ice nuclei would form within the interior volume of the droplet. How-
4 ever Tabazadeh et al. (2002b) proposed that the gas-liquid interface may be
5 important for homogeneous freezing of atmospheric droplets. The relative im-
6 portance of surface- and volume-based nucleation within their scheme depends
7 on the temperature and size of the droplet. Smaller droplets are more likely
8 to undergo surface nucleation, while lowering the temperature increases the
9 likelihood of volume-based nucleation occurring. A recent laboratory study
10 (Duft and Leisner, 2004) has demonstrated that homogeneous freezing in wa-
11 ter droplets with radii 19–49 μm is volume-based, although further study is
12 required to discover if this is true for smaller droplets. Computational simula-
13 tions (Vrbka and Jungwirth, 2006) suggest that, during homogeneous freezing,
14 ice nuclei are more likely to form within the sub-surface region. They suggest
15 this is because the change in volume associated with freezing is more easily
16 accommodated near the surface than in the bulk of the liquid. However, the
17 disorder of the immediate liquid surface is not conducive to crystal formation.
18 Liquid PSC particles are not pure water, so the water droplet simulations are
19 not directly applicable to PSCs, and it has been noted that the surface of an
20 aqueous particle is often enriched with chemical components such as nitric
21 acid, making it more difficult for an ice nucleus to form at the surface of such
22 a particle (see Tabazadeh et al., 2002b, and references within).

23 Theoretical work on heterogeneous nucleation of ice particles from aqueous
24 solutions in the troposphere suggests that this process could be important for
25 the formation of cirrus clouds in the northern hemisphere (e.g., see Gierens,
26 2003, and references within). However, the total abundance of condensation

1 nuclei is far less in the lower stratosphere, $10\text{--}60\text{ cm}^{-3}$, than in the upper tro-
2 posphere, $600\text{--}1600\text{ cm}^{-3}$ (Sheridan et al., 1994), and it is believed that only
3 a fraction of these are likely to act as ice nuclei. From this we might infer
4 that heterogeneous ice nucleation is less relevant in the stratosphere, although
5 recent laboratory experiments have shown that fumed silica can act as a nu-
6 cleus for the heterogeneous freezing of ice from dilute ($\sim 7\%$ HNO_3) nitric acid
7 solutions (Bogdan et al., 2003). Since fumed silica is similar in composition,
8 and method of formation, to meteoritic smoke particles, then these particles,
9 which have a number concentration of 100 cm^{-3} in the stratosphere, could act
10 as heterogeneous nuclei for ice particles (see Bogdan et al., 2003, and references
11 within).

12 The equilibrium partitioning of H_2O between the solid and gas phases for a
13 water ice particle can be calculated using, for example, the empirical equation
14 derived by Marti and Mauersberger (1993). STS particles are, however, not
15 composed purely of water, and so the ice freezing process must take account
16 of the other chemical constituents of the droplet. One possibility is that the
17 other components of the droplet, in this case mostly HNO_3 and H_2SO_4 , could
18 simply be adsorbed onto the surface of the ice particle as it grows, and become
19 embedded in the ice matrix without forming hydrates. However the thermo-
20 dynamic solubilities of HNO_3 and H_2SO_4 in ice are of the order of parts per
21 million (Davy and Somorjai, 1971), at least five orders of magnitude less than
22 their mixing ratios in PSCs. Instead it is likely that the remaining part of the
23 original STS will form either solid NAT/SAT or a liquid STS layer on the
24 surface of the ice particle.

25 Using the thermodynamic model of Carslaw et al. (1995), Koop et al. (1997a)
26 investigated the stabilities of the different possible PSC phases. They found

1 that liquid solutions can co-exist with both NAT and ice, provided that H_2SO_4
2 remains in the aqueous phase. Liquids in equilibrium with a solid phase con-
3 sisting just of ice contain ~ 40 wt% HNO_3 and trace H_2SO_4 , whilst equilib-
4 rium with NAT would result in a H_2SO_4 -rich solution. Liquids in equilibrium
5 with SAT would have HNO_3 vapour pressures in excess of the total available
6 HNO_3 , and so are thermodynamically unfeasible. So a Type II particle is likely
7 to consist of a combination of different chemical phases, with typically combi-
8 nations being: ice/STS, ice/NAT/STS (almost binary $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ solution),
9 and ice/NAT/SAT.

10 3.3 *Type Ia particles*

11 Type Ia particles are classified as medium-sized ($\geq 1 \mu\text{m}$) aspherical parti-
12 cles, which can exist several degrees above the ice frost point (Toon et al.,
13 1990). Their aspherical shape implies that they are solid particles, most likely
14 nitric-acid hydrates. Experimental work by Hanson and Mauersberger (1988)
15 demonstrated that the stable nitric-acid hydrate under stratospheric condi-
16 tions is nitric acid trihydrate (NAT, $\text{HNO}_3 \cdot 3\text{H}_2\text{O}$, ~ 54 wt% HNO_3), which
17 can exist up to about 7 K above the ice frost point. Later experiments by
18 Worsnop et al. (1993) showed that other nitric-acid hydrates are metastable
19 under stratospheric conditions (Figure 11). They suggested that nitric acid
20 dihydrate (NAD, $\text{HNO}_3 \cdot 2\text{H}_2\text{O}$, ~ 64 wt% HNO_3) would form readily in the
21 stratosphere and could persist for several days; a conclusion which is sup-
22 ported by recent studies into NAD crystalline structures and formation routes
23 (Natsheh et al., 2006). This tendency to nucleate metastable phases may in-
24 dicate a relatively high free-energy barrier for nucleation of NAT compared

1 with other crystalline nitric-acid hydrates (Worsnop et al., 1993). Although
2 the existence of nitric-acid hydrates in the polar winter stratosphere has been
3 known to be possible since the mid-1980s, it has only been recently that direct
4 in-situ measurements of particle compositions have confirmed the existence of
5 NAT particles in the stratosphere (Voigt et al., 2000a). No direct stratospheric
6 measurements of NAD particles have been made.

7 As with stratospheric ice particles, there have been three routes of formation
8 suggested for HNO_3 hydrates: vapour deposition onto either ice or SAT, ho-
9 mogeneous nucleation from an aqueous solution, and heterogeneous nucleation
10 in aqueous solutions containing solid cores. The original three-stage model of
11 PSC formation implied that NAT would nucleate directly from the gas-phase
12 on to SAT particles. However, theoretical studies cast doubt on this mech-
13 anism (MacKenzie et al., 1995), and laboratory studies (Iraci et al., 1995)
14 showed that the formation of NAT on SAT is strongly hindered up to NAT
15 saturation ratios of 127, implying that this process is kinetically very unlikely
16 in the stratosphere. A possible exception to this is when SAT surfaces have
17 been primed for NAT nucleation by a previous deposition-and-evaporation cy-
18 cle of NAT on the surface (Zhang et al., 1996). This process could occur in
19 the stratosphere for SAT particles which originally formed as part of a multi-
20 phase particle which included NAT, but the formation route for these particles
21 is uncertain. In addition, the NAT saturation ratios needed for deposition onto
22 SAT occur below the deliquescence temperature of SAT in the stratosphere.

23 Bulk freezing experiments (Koop et al., 1997b) showed that the nucleation of
24 NAT from binary $\text{HNO}_3/\text{H}_2\text{O}$ solutions containing ice particles occurred faster
25 if the ice particles had passed through the gas-phase first. Koop et al. (1997b)
26 interpreted this to mean that NAT/NAD had nucleated heterogeneously from

1 the gas phase on to the ice particles while they were sedimenting. When the
2 crystals reached the liquid they seeded NAT and ice formation in the binary
3 solution. The nucleation of NAT by vapour deposition onto ice surfaces has
4 been observed at high NAT supersaturations (Biermann et al., 1998; Middle-
5 brook et al., 1996). Modelling studies (Luo et al., 2003) have suggested that
6 rapid cooling within gravity waves would suppress the growth of STS parti-
7 cles, allowing the development of large NAT saturation ratios in parallel with
8 the growth of ice particles (Figure 12). This could result in the nucleation of
9 NAT onto the ice particles, which would then persist as NAT particles after
10 the ice had evaporated away (Luo et al., 2003).

11 The most direct method of NAT formation would be homogeneous freezing
12 from STS droplets. However there are severe limitations. Bulk freezing experi-
13 ments (Koop et al., 1995, 1997b) demonstrated that the presence of ≥ 0.1 wt%
14 H_2SO_4 would lower the freezing temperatures of $\text{HNO}_3/\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ solutions
15 by ≥ 10 K below those typical of $\text{HNO}_3/\text{H}_2\text{O}$ solutions. Bertram et al. (1996);
16 Disselkamp et al. (1996), and Bertram and Sloan (1998) found somewhat dif-
17 ferent results while studying the freezing of 1:2 ($\text{HNO}_3:\text{H}_2\text{O}$ molar ratio) nitric
18 acid aerosol droplets to NAD. Freezing in the experiments of Disselkamp et al.
19 (1996) produced NAD much more easily from solutions with a $\text{HNO}_3:\text{H}_2\text{O}$
20 molar ratio of 1:2 than NAT from solutions with a molar ratio of 1:3 (the
21 nucleation rates differ by a factor of 10^3 or more). Freezing temperatures in
22 the experiments of Bertram et al. (1996); Bertram and Sloan (1998) were
23 some 20 K below those in the Disselkamp et al. (1996) experiments. It is not
24 clear why these experimental results are so different. Classical freezing theory
25 — that is also consistent with NAT, SAT, and ice freezing experiments —
26 is consistent with the Bertram et al. (1996); Bertram and Sloan (1998) re-

1 sults (MacKenzie et al., 1998). Subsequent empirical fits to laboratory results
2 (Möhler et al., 2006) fit the results of the most recent study (Stetzer et al.,
3 2006) and those of the earlier studies, using a fitted version of classical freezing
4 theory.

5 As discussed previously, Meilinger et al. (1995) suggested that these limitations
6 might be overcome during the rapid changes in conditions which occur within
7 gravity waves. However the maximum HNO_3 mass fraction they reported was
8 $\sim 52\%$ in cooling rates of up to 36 K hr^{-1} , which is close to the NAT stoi-
9 chiometry, but not that of NAD. Tsias et al. (1997) expanded on this work,
10 examining the effect of faster temperature changes on the droplet composi-
11 tions. They found that particles could reach HNO_3 mass fractions of $\sim 58\%$
12 ($\sim 1:2.5 \text{ HNO}_3/\text{H}_2\text{O}$) during warming at rates of $\geq 100 \text{ K hr}^{-1}$ (temperature
13 change rates in stratospheric mountain wave events can be up to 150 K hr^{-1} ,
14 e.g. see Tsias et al., 1997, and references within). They demonstrated ex-
15 perimentally that binary $\text{HNO}_3/\text{H}_2\text{O}$ droplets of similar compositions would
16 freeze to NAD typically within one hour. During mountain wave events, how-
17 ever, such compositions are only achieved for a few minutes, so only a small
18 fraction of the droplets at such compositions (which would be a small fraction
19 of the total aerosol population) would freeze.

20 Experimental studies of binary $\text{HNO}_3/\text{H}_2\text{O}$ droplet freezing (Salcedo et al.,
21 2000, 2001) were extrapolated linearly to stratospheric conditions and used
22 to hypothesise a “nucleation window” for the production of NAD (Tabazadeh
23 et al., 2001). Such a linear extrapolation is forbidden by thermodynamics
24 (Knopf et al., 2002) and particle production rates (for NAT and NAD) are
25 much smaller than those predicted by Tabazadeh et al. (2001) (Figure 13b).

1 While all of the studies above have assumed that homogeneous nucleation
2 would occur within the bulk of a liquid, other work (Tabazadeh et al., 2002a,b)
3 has proposed surface-based nucleation. In this process, the nucleating solid
4 reduces the energy needed to form a stable crystal by forming part of its
5 surface out of the liquid, in contact with the gas-phase. This method is only
6 favoured in cases where the solid is partially (but not fully) wettable by its
7 own liquid melt. Tabazadeh et al. (2002a) re-analysed data from laboratory
8 studies of the freezing of NAT and NAD from binary $\text{HNO}_3/\text{H}_2\text{O}$ solutions
9 using a formulation for surface-based nucleation. They found that the for-
10 mulation for surface nucleation fitted the data better than the volume-based
11 formulation and is, at least in the case of NAD nucleation, less temperature
12 dependent. A comparison of atmospheric NAD production rates derived using
13 the two formulations (Figure 14), show that the surface-based production rate
14 is $\sim 10^2$ greater than the volume-based production rate used in their earlier
15 work (Tabazadeh et al., 2001). Applying the factor of 10^2 to the stratospheric
16 production rate of NAD of Knopf et al. (2002) (as shown in Figure 13b) we
17 only obtain a production rate of $\sim 10^{-11} \text{ cm}^{-3} \text{ hr}^{-1}$. This is still much smaller
18 than the production rates of $10^{-5} \text{ cm}^{-3} \text{ hr}^{-1}$ that microphysical sensitivity
19 studies suggest are needed for homogeneous nucleation to have a significant
20 influence on the overall number of HNO_3 -hydrate particles in the stratosphere
21 (Tabazadeh et al., 2001). Recently Stetzer et al. (2006) have demonstrated
22 that their measurements of freezing rates of NAD from small ($0.25 \mu\text{m}$) binary
23 HNO_3 droplets, in the low temperature aerosol chamber AIDA, agree well
24 with the results of a number other freezing studies in the literature when us-
25 ing volume-based nucleation rates. The surface-based nucleation rates Stetzer
26 et al. (2006) obtained from their data were also 1–3 orders of magnitude lower
27 than those calculated by Tabazadeh et al. (2002a). They concluded that, for

1 binary HNO_3 droplets, volume-based, rather than surface-based, homogeneous
2 nucleation is more likely under stratospheric conditions. Their companion pa-
3 per (Möhler et al., 2006) estimates very small homogeneous (volume-based)
4 nucleation rates for NAT in the stratosphere.

5 When considering liquid particle microphysics, we need to take into account
6 the surface enrichment of HNO_3 in both binary $\text{HNO}_3/\text{H}_2\text{O}$ and ternary
7 $\text{HNO}_3/\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ solutions (Yang and Finlayson-Pitts, 2001; Donaldson and
8 Anderson, 1999), which could influence surface-based nucleation rates. In addi-
9 tion, the presence of H_2SO_4 in STS droplets can strongly inhibit the nucleation
10 of HNO_3 -hydrates (Koop et al., 1997b). The Koop et al. (1997b) experiments,
11 however, used bulk samples, which are not suitable for studying surface-based
12 nucleation. Tabazadeh et al. (2002a) have suggested that, if H_2SO_4 is not a
13 “surface active” component, then it will have little effect on the surface tension
14 of the solution, and so on the free energy barrier to nucleation. However more
15 measurements of the nucleation of NAT/NAD in ternary solution droplets, at
16 stratospherically relevant temperatures, are needed in order to answer these
17 questions. In conclusion it appears that, from current knowledge, homogeneous
18 nucleation of NAT/NAD from STS is not significant in the stratosphere, al-
19 though there is not universal consensus on this subject.

20 As mentioned above, bulk freezing experiments (Koop et al., 1997b) have
21 shown that heterogeneous nucleation of NAT from nitric-acid solutions on to
22 ice is too slow a process to be significant in the stratosphere. In addition, lab-
23 oratory measurements (Zondlo et al., 2000) indicate that binary $\text{HNO}_3/\text{H}_2\text{O}$
24 solutions on an ice surface do not nucleate to NAT readily, but form a meta-
25 stable liquid $\text{HNO}_3/\text{H}_2\text{O}$ film instead.

1 Laboratory experiments using finely divided aqueous systems (Bogdan et al.,
2 2003) have shown that binary $\text{HNO}_3/\text{H}_2\text{O}$ solutions with ~ 53 wt% HNO_3
3 will heterogeneously freeze to nitric acid hydrate on fumed silica at around
4 195 K. These results suggest that heterogeneous nucleation of NAT/NAD on
5 meteoritic smoke nuclei could occur on STS particles passing through gravity
6 waves, in a similar manner to the non-equilibrium homogeneous freezing de-
7 scribed above. Similar experiments with binary $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ solutions found
8 that freezing did not occur until around 175–180 K, too low a temperature for
9 the stratosphere. This suggests that, again in a similar manner to homogeneous
10 freezing of STS (Koop et al., 1997b), the heterogeneous freezing of liquid PSCs
11 will only occur when H_2SO_4 is a minor constituent of the droplets (≤ 0.1 wt%),
12 although only freezing experiments with ternary $\text{HNO}_3/\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ solutions
13 will prove this theory.

14 **4 The effects of PSCs**

15 Soon after the discovery of the ozone hole PSCs were found to be important
16 both for aiding the release of active chlorine species as well extending the
17 lifetimes of these species by removing HNO_3 (Toon et al., 1986; Solomon et al.,
18 1986; Solomon, 1990, 1999).

19 Below we review how the microphysical properties of the PSCs determine the
20 magnitude and extent of these effects, and how the microphysical properties
21 are, in turn, dependent on the climate of the polar vortex. The Antarctic
22 vortex is generally more stable, with greater symmetry around the pole, than
23 the Arctic vortex. This is due to the more disruptive atmospheric wave activity
24 in the Northern hemisphere, which is caused by the configuration of land and

1 sea in that hemisphere (see, e.g., Solomon, 2004, for a succinct description). As
2 a result, the temperature distribution, and so volume of the air in which PSCs
3 can exist, of the Arctic vortex depends greatly on the dynamical situation
4 of each winter (Pawson and Naujokat, 1999). As a consequence chlorine is
5 frequently fully activated within the Antarctic vortex and so ozone loss is large
6 (e.g. WMO, 2006). Ozone loss within the Arctic vortex can vary considerably
7 from year to year, and is very dependent on changes in climate conditions
8 (Rex et al., 2004). The interactions between stratospheric ozone and climate
9 are outside the scope of this review — the interested reader is directed to
10 WMO (2006); Eyring et al. (2005) and references therein — but we note in
11 passing that global warming at the surface is associated with cooling in the
12 stratosphere. This is because greenhouse gases trap terrestrial infrared near the
13 Earth’s surface and reduce the flux entering the stratosphere. There is evidence
14 that the volume of Arctic stratosphere susceptible to ozone destruction may
15 be changing, particularly in those years that are generally cold (Rex et al.,
16 2004).

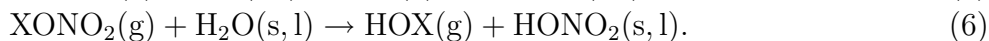
17 Lidar observations of PSCs from Ny-Ålesund and McMurdo are illustrative,
18 though not representative, of the differences in current PSC climatology be-
19 tween the two hemispheres (Maturilli et al., 2005). The number of days with
20 PSC observations was found to be 5 times higher at McMurdo (78°S, 167°E)
21 than at Ny-Ålesund (79°N, 12°E). PSCs above McMurdo generally contained
22 NAT, with very few observations of liquid or ice particles. In contrast, liq-
23 uid PSCs were more common than NAT PSCs above Ny-Ålesund, however
24 almost two-thirds of the days on which PSCs were observed both liquid and
25 NAT PSCs were present.

26 Below we will examine the current knowledge of how the microphysics of PSC

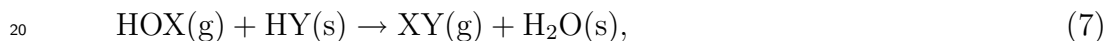
1 particles affect heterogeneous reactions, and their role in removing HNO_3 . We
2 do not undertake a discussion of stratospheric chemistry in general; this can
3 be found in, e.g. Brasseur and Solomon (2005); Seinfeld and Pandis (2006);
4 MacKenzie (2003). Detailed discussions of outstanding problems in strato-
5 spheric chemistry are given in the special issue of *Chemical Reviews* on long-
6 term issues in atmospheric chemistry (Ravishankara, 2003).

7 4.1 Heterogeneous Reactions on PSC particles

8 Investigation into the properties of PSCs has been driven by their role in the
9 destruction of the ozone layer. It was Farman et al. (1985) who first high-
10 lighted the thinning of the Southern polar ozone layer during the austral
11 spring. Solomon et al. (1986) demonstrated that ozone depletion was confined
12 to the lower polar stratosphere (10–20 km), which McCormick et al. (1981)
13 had shown to contain PSCs during the austral winter. Solomon et al. (1986)
14 proposed two new heterogeneous reactions which, taking place on the surfaces
15 of PSC particles, would release active halogen species, principally Cl_2 , from
16 inert reservoir species:



17 where X and Y can be either Cl or Br. The photolytically active gaseous
18 molecules, XY and HOX, resulting from reactions 5 and 6, as well as the
19 reaction



1 would release atomic chlorine or bromine with the onset of spring, initiating
2 the catalytic destruction of ozone (Solomon, 1999).

3 The rates of reactions 5–7 due to the interaction of gas and aerosol can be
4 represented by the molecular flux J

$$5 \quad J = \gamma n_g \bar{c} / 4, \quad (8)$$

6 where γ is the uptake coefficient, n_g is the density of the gas species and \bar{c}
7 is the mean gas molecular speed. J depends on a number of processes: gas-
8 phase diffusion, mass accommodation, Henry’s law solubility (in the case of
9 liquids), chemical reactions in the condensed bulk phase, and particle surface
10 properties. These processes are coupled to each other, making their individual
11 determination from laboratory data difficult. Instead, in reactive uptake cases,
12 γ can be approximated by

$$13 \quad \frac{1}{\gamma} = \frac{1}{\alpha} + \frac{1}{\gamma_{sol} + \gamma_{rxn}}, \quad (9)$$

14 where α is the mass accommodation coefficient for the relevant species, γ_{sol}
15 is the solubility uptake coefficient, and γ_{rxn} is the reactive uptake coefficient.
16 In cases of low solubility, or long exposure time, the system reaches a steady
17 state and γ_{sol} becomes negligible; γ comes to be governed by α and γ_{rxn} .

18 Uptake coefficients for many of the different permutations of reactions 5–7 on
19 typical stratospheric aerosol types have been derived using laboratory data.
20 State of the art values for these are documented in the reviews published by
21 JPL and IUPAC (respectively available from <http://jpldataeval.jpl.nasa.gov>
22 and <http://www.iupac-kinetic.ch.cam.ac.uk>). Experimental results indicate that
23 uptake coefficients for most combinations of reactions 5–7 on water ice are

1 less than 0.3 (Sander et al., 2006). Experimental results indicate similarly
2 small uptake coefficients for reactions on nitric acid ice, although parameter-
3 isations are only available for the reactions $\text{ClONO}_2 + \text{H}_2\text{O}$, $\text{ClONO}_2 + \text{HX}$,
4 and $\text{HOCl} + \text{HCl}$. Studies of reactions 5–7 on sulphuric acid solutions have
5 shown a strong negative temperature dependence — caused by the dependence
6 of these reactions on the solubility of the halogen species and on the activity of
7 H_2O , which have inverse relationships with temperature. The most recent, and
8 comprehensive, parameterisation of reactions 5–7 for chlorine species is given
9 in Shi et al. (2001). Looking at their parameterisation we see that the uptake
10 coefficients for all the reactions increase as the temperature decreases (and
11 the H_2SO_4 mass fraction drops), until just under 50% H_2SO_4 , when reaction 5
12 wins out in the competition with the other two reactions and approaches unity,
13 while the other reactions don't get above 0.1 (Figure 15). Reactions 5 and 6
14 have been parameterised for chlorine over SAT (Zhang et al., 1994a), how-
15 ever no more parameterisations are available for reactions occurring on SAT,
16 likewise for reactions occurring on STS. Instead, for STS droplets, current
17 microphysical models use the reaction rates for binary sulphuric acid droplets
18 (e.g. Drdla and Schoeberl, 2003).

19 The activation rates of halogens on aerosol particles are influenced by both
20 the uptake coefficient on the particle and the particulate surface area avail-
21 able. Although we do not have measured values for all these reactions on all
22 possible stratospheric aerosol types, we can see that, for the temperatures at
23 which PSCs exist, the uptake coefficients are much greater for the background
24 stratospheric liquid sulphate aerosols than for the solid PSC particles. The
25 available surface area of the liquid sulphate aerosol is generally much higher
26 than that of the solid PSCs – the mean background stratospheric aerosol sur-

1 face area density is $1\mu\text{m}^2\text{cm}^{-3}$ (e.g. Thomason et al., 1997) – but very much
2 less than the surface area density of liquid PSCs. Within the stable Antarctic
3 vortex, NAT PSCs with relatively large surface area densities are widespread;
4 in the Arctic vortex, however, such dense NAT PSCs tend to be limited to the
5 relatively small cold regions induced by mountain waves (e.g. Carslaw et al.,
6 1998b; Toon et al., 2000). Water ice PSCs are even less common, due to their
7 low formation temperatures, however when they do form they can provide
8 aerosol surface areas 100 times greater than those of liquid or NAT PSCs (e.g.
9 Carslaw et al., 1998b).

10 Since the onset of chlorine activation occurs at ~ 195 K (Toohey et al., 1993;
11 Webster et al., 1993), it was initially assumed that activation was due to type
12 I (i.e. NAT) PSCs. As knowledge of PSC microphysics has improved it has be-
13 come clear that chlorine activation is due more to the presence of liquid rather
14 than NAT particles. However, as a first order approximation, many large-scale
15 simulations of stratospheric ozone loss still use T_{NAT} or 195 K as the thresh-
16 old temperature for switching on heterogeneous chlorine activation (e.g. Kühl
17 et al., 2004; Geer et al., 2006). Drdla (2007) suggests that the coincidence of
18 chlorine activation with T_{NAT} occurs because at standard stratospheric con-
19 ditions this also happens to be close to the temperature at which chlorine
20 activation becomes significant on the background liquid sulphate aerosols.
21 However this relationship can break down in air masses which have under-
22 gone perturbations such as denitrification. Drdla (2007) have developed a pa-
23 rameterisation (called T_{ACL}) — based on H_2O mixing ratio, sulphate surface
24 area at 210 K, and pressure — which can be used as a diagnostic of where chlo-
25 rine activation is most likely. Finding a sufficiently accurate and computation-
26 ally efficient parameterisation of PSCs and chlorine activation for global-scale

1 models is one of the most important remaining problems in PSC studies.

2 As noted above, current microphysical models which deal with chlorine ac-
3 tivation on STS droplets must use reaction rates devised for sulphuric acid
4 solutions. Measurements of reactions 5–7 on ternary have been made, but
5 no clear parameterisation has been produced. Zhang et al. (1994b) measured
6 reactions 6 and 7, and concluded that the rates of these reactions were un-
7 affected by the presence of HNO_3 . However none of their solutions contained
8 more than 5% HNO_3 . Elrod et al. (1995) measured reaction 6 over ternary so-
9 lutions with $\text{HNO}_3/\text{H}_2\text{SO}_4$ mass fractions between 0.283/0.202 – 0.022/0.503.
10 The reaction probabilities they measured for these solutions were between
11 5–32% lower than those they measured for sulphuric acid solutions at repre-
12 sentative compositions for the relevant stratospheric temperatures. However
13 there was no obvious dependence on composition in these reductions. Zhang
14 et al. (1995) measured Reaction 5 on ternary solution over the temperature
15 range 196–214 K. The $\text{HNO}_3/\text{H}_2\text{SO}_4$ mass fractions at 196 K were 0.112/0.364
16 (as calculated by Hanson, 1998). They concluded that the probabilities of
17 Reaction 5 were the same for ternary solutions as they were for sulphuric
18 acid solutions; however Hanson (1998) noted that, where appreciable amounts
19 of HNO_3 are present, their data are consistent (within uncertainties) with a
20 significant reduction in the reaction probabilities. Hanson (1998) measured
21 Reactions 5 and 6 on ternary solutions with $\text{HNO}_3/\text{H}_2\text{SO}_4$ mass fractions of
22 0.044/0.446 to 0.256/0.203. They found that the measured reaction proba-
23 bilities were significantly lower than those predicted for the corresponding
24 sulphuric acid solutions. Fits to the data were obtained by reducing the bulk
25 reaction probability by a factor of 2, and the surface reaction probabilities by a
26 factor of 0–10 (increasing with increasing HNO_3 mass fraction). However they

1 produced no comprehensive parameterisation of the reaction probabilities on
2 ternary solutions.

3 To date there have been no more experimental investigations of the rates of
4 reactions 5–7 on ternary solutions. The lack of data on these reactions is ham-
5 pering our understanding of chlorine activation in the Arctic vortex, where
6 STS PSCs are very common (e.g. Maturilli et al., 2005). The increases in
7 particle surface area of STS aerosols over sulphuric acid aerosols, due to the
8 uptake in nitric acid and water, will lead to a greater activation of chlorine.
9 However this increase may not be as large as that currently seen in microphys-
10 ical models if the presence of HNO_3 reduces the reaction rates by the degree
11 seen in the studies of Elrod et al. (1995) and Hanson (1998)(Lowe et al., 2006).

12 4.2 Denitrification

13 In addition to providing reactive surfaces for the release of active chlorine
14 compounds, PSCs can contribute to the destruction of ozone via the denitri-
15 fication of the stratosphere. The condensation of HNO_3 on to PSC particles
16 stops the deactivation of gas-phase ClO_x by reaction with NO_2 from photolysed
17 HNO_3 (Toon et al., 1986). If PSC particles subsequently sediment out of the
18 stratosphere, then the HNO_3 they contain is permanently removed, extending
19 the lifetime of the active chlorine compounds. Denitrification occurs in both
20 polar stratospheres (e.g. Fahey et al., 1990), however it is most intense over
21 Antarctica; where large fractions of the total HNO_3 are irreversibly removed.

22 The differences in the amounts of denitrification, and ozone loss, between the
23 two hemispheres are due to the greater stability of the Antarctic vortex. Al-

1 though both vortices form at similar times in the autumn/early-winter, on
2 average the Antarctic vortex breaks down a month later in the season than
3 the Arctic vortex. In addition, while the Antarctic vortex is generally very
4 stable except during the spring breakup; the area, position, and strength of
5 the Arctic vortex are highly variable, especially during the late winter (Waugh
6 and Randel, 1999). The stability of the Antarctic vortex produces much colder
7 conditions than are experienced in the Arctic stratosphere, allowing increased
8 PSC formation which leads to greater denitrification. In addition, extend-
9 ing the cold stratospheric temperatures into the austral spring allows the re-
10 activation, on PSCs, of any ClONO_2 formed during the spring. Portmann et al.
11 (1996) argue that this process is as important as denitrification for ozone loss
12 in the Antarctic.

13 Denitrification involves the sedimentation of large HNO_3 -containing particles;
14 however, until recently, it has not been clear exactly what these particles are,
15 or how they form. STS particles are too small (typically radii of $\sim 0.5 \mu\text{m}$) to
16 have sedimentation speeds greater than a metre per day. Similarly, if all back-
17 ground aerosol particles act as nucleation sites for NAT, then NAT particles
18 would also be too small to promote denitrification. The most likely particles to
19 cause denitrification are ice particles. Because of the relatively high amounts of
20 H_2O in the atmosphere, ice particles can grow to the large sizes necessary for
21 efficient sedimentation, and can incorporate gaseous HNO_3 in the form of NAT
22 (e.g. Wofsy et al., 1990). The sedimentation of these ice particles would also
23 result in dehydration of the stratosphere, which inhibits ozone loss, though to
24 a lesser degree than it is promoted by denitrification.

25 Observations of denitrification have, however, painted a more complicated
26 picture. Analysis of in-situ and remote observations from the Arctic winter of

1 1994–95 suggested that the denitrification observed was predominantly caused
2 by small number densities of large NAT particles (Waibel et al., 1999). The
3 then-current modelling practice of assuming that all aerosol particles froze to
4 NAT, creating a population of many small NAT particles, led to an under-
5 estimation of the true influence of NAT particles on the stratosphere. In addi-
6 tion, satellite-based observations indicated that, during relatively warm winter
7 months in the Antarctic, extensive denitrification can occur without any sig-
8 nificant dehydration (Tabazadeh et al., 2000). However, up to this point, no
9 large NAT particles had been measured in-situ.

10 This changed in the Arctic winter of 1999–2000, during which the most exten-
11 sive denitrification observed above the Arctic occurred (Popp et al., 2001). In-
12 situ measurements (see Section 2 and Northway et al., 2002) detected HNO₃-
13 containing particles, with diameters as large as 10–20 μm and total number
14 densities in the order of 10^{-4} cm^{-3} (Fahey et al., 2001). Highly selective nucle-
15 ation mechanisms, such as low freezing rates of NAT (or NAD) below $\sim 196 \text{ K}$
16 (e.g. Tabazadeh et al., 2001), or NAT formation via ice particles (e.g. Waibel
17 et al., 1999), were put forward to interpret the observations. Because of the
18 length of time needed for these particles to grow, Mann et al. (2003) showed
19 that, regardless of nucleation mechanism, the amount of denitrification likely
20 to occur could be determined using the “closed-flow area” of the vortex. This
21 is based on the degree of concentricity of the vortex and the mass of air with
22 temperatures below T_{NAT} (see Figure 16).

23 Model studies (Carslaw et al., 2002) have shown that synoptic-scale ice clouds
24 were very unlikely to be the source of the observed “NAT-rocks”, and sug-
25 gested that, although localised gravity waves are known sources of NAT parti-
26 cles (Carslaw et al., 1998a), they would produce only a limited number of the

1 observed large particles. Fueglistaler et al. (2002) proposed that the gravita-
2 tional settling of particles from typical Type Ia PSCs, into air supersaturated
3 with respect to NAT, could act as a source for NAT-rocks. However observa-
4 tions of NAT PSCs in situations where this cannot have occurred suggest that
5 this mechanism is not the only source of NAT-rock particles in the stratosphere
6 (Voigt et al., 2005). A number of studies of NAT formation during the win-
7 ter of 2002/2003 have suggested that observed NAT particle numbers (Larsen
8 et al., 2004; Voigt et al., 2005) and levels of denitrification (Grooß et al., 2005;
9 Davies et al., 2006) can be reproduced using a simple flat nucleation rate of
10 NAT particles, at temperatures below T_{NAT} . The nucleation rates determined
11 are 25×10^{-6} (Larsen et al., 2004), 8×10^{-6} (Voigt et al., 2005), 7.8×10^{-6}
12 (Grooß et al., 2005), and 12×10^{-6} particles $\text{cm}^{-3}\text{hr}^{-1}$ (Davies et al., 2006).
13 Modelling studies of denitrification during previous Arctic winters have de-
14 termined optimal nucleation rates of a similar magnitude: 2.9×10^{-6} parti-
15 cles $\text{cm}^{-3}\text{hr}^{-1}$ for the winter 1999/2000 (Mann et al., 2003), and 12×10^{-6}
16 particles $\text{cm}^{-3}\text{hr}^{-1}$ for the winter 1994/95 (Davies et al., 2005).

17 **5 Summary**

18 To summarise: **(1)** In-situ instrumentation has improved to the point where
19 we can measure the amounts of the major constituents of the condensed phase
20 and their ratios — allowing for the positive identification of STS and NAT
21 particles in the Arctic Stratosphere. **(2)** Satellite instruments can now be used
22 to identify the category of PSCs as well as their synoptic extent - improving
23 the temporal and spatial coverage of observations. **(3)** The major types of PSC
24 particles have been identified, and observed in the stratosphere. Simple ho-

1 homogeneous nucleation of NAT from STS droplets has been shown to occur at
2 too low a rate to be significant in the stratosphere. Whether NAT nucleates
3 through a classical volume-dependent heterogeneous route, a surface-based
4 pseudo-heterogeneous route, or some other route, is still not clear. **(4)** Our
5 knowledge of the heterogeneous reaction rates for the release of halogens on
6 PSC particles is extensive — but still inadequate for STS droplets, which
7 are the most common PSC type in the Arctic. This lack of knowledge lim-
8 its the accuracy of the large-scale microphysical models which have recently
9 been developed. **(5)** “NAT-rocks” have been discovered in the Arctic. These
10 large particles have been shown to be important in the denitrification of the
11 stratosphere. **(6)** Global warming at the surface is associated with cooling in
12 the stratosphere. Stratospheric cooling will likely change the frequency of oc-
13 currence of PSCs, as well as changing stratospheric chemistry more generally.
14 There is evidence that the volume of Arctic stratosphere susceptible to ozone
15 destruction may be changing, particularly in those years that are generally
16 cold. **(7)** A sufficiently accurate, yet computationally efficient, parameterisa-
17 tion of PSC microphysics and chemistry, for use in global chemistry-climate
18 models, is yet to be developed. This limits the accuracy with which the recov-
19 ery of the ozone layer, and the impact of stratospheric processes on climate,
20 can be forecast.

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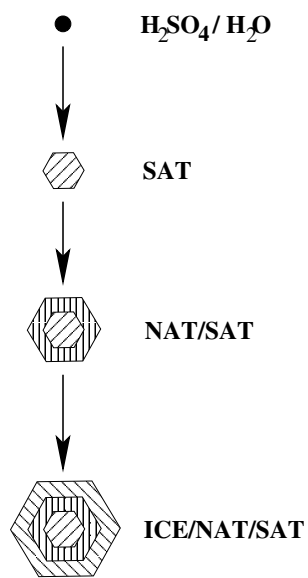


Fig. 1. The 3-stage model of PSC development.

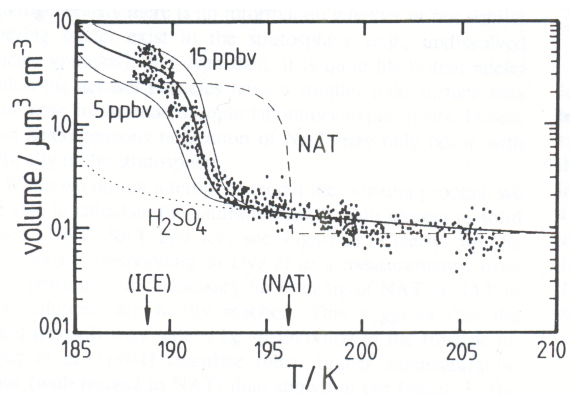


Fig. 2. Particle volumes of Dye et al. (1992) compared with model calculations. From Carslaw et al. (1994). Copyright 1994 American Geophysical Union. Reproduced by permission of American Geophysical Union.

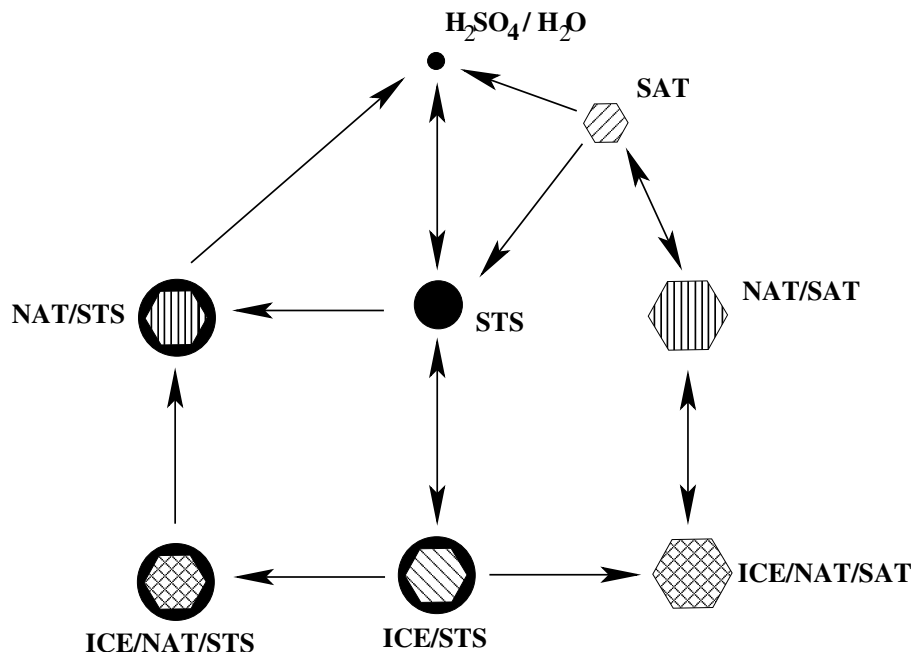


Fig. 3. The STS model of PSC development. Laboratory studies, field observations, and theory suggest that some routes between particle types are bi-directional, while others are only one-way. The NAT/STS and NAT/SAT phases encompass all ‘solid’ PSC types, including NAT rocks.

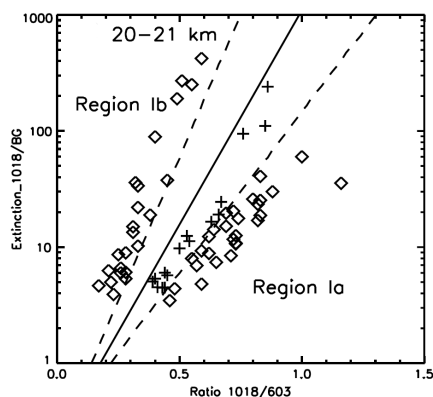


Fig. 4. POAM observations of PSCs during the Arctic winter of 1999/2000 between 20 and 21 km altitude. For each observation the wavelength dependence of the extinction coefficient (shown as the ratio of the $1.018 \mu\text{m}$ to $0.603 \mu\text{m}$ extinctions) is plotted against the $1.018 \mu\text{m}$ extinction coefficient (which is normalised using the background aerosol extinction coefficient). The observations classified as either Type Ia or Ib are shown by the diamonds (in the relevant region); the crosses represent undetermined observations. The solid line represents the division between the criteria for the two different types of cloud, while the dashed lines represent the uncertainty envelope for this division. From Strawa et al. (2002). Copyright 2002 American Geophysical Union. Reproduced by permission of American Geophysical Union.

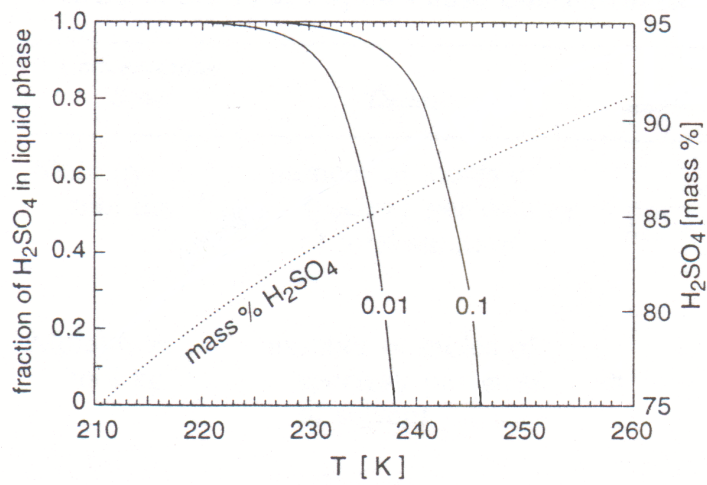


Fig. 5. Phase of sulphuric acid at 10 mbar (~ 30 km) altitude against temperature, assuming 5 ppmv water vapour. The solid lines indicate the fraction of the total H_2SO_4 (left-hand scale) present in the condensed phase for 0.01 and 0.1 ppbv total H_2SO_4 . The dotted line indicates shows the mass fraction of H_2SO_4 (right-hand scale) in binary $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ droplets in equilibrium with the specified water vapour mixing ratio and temperature. Adapted from Carslaw et al. (1997).

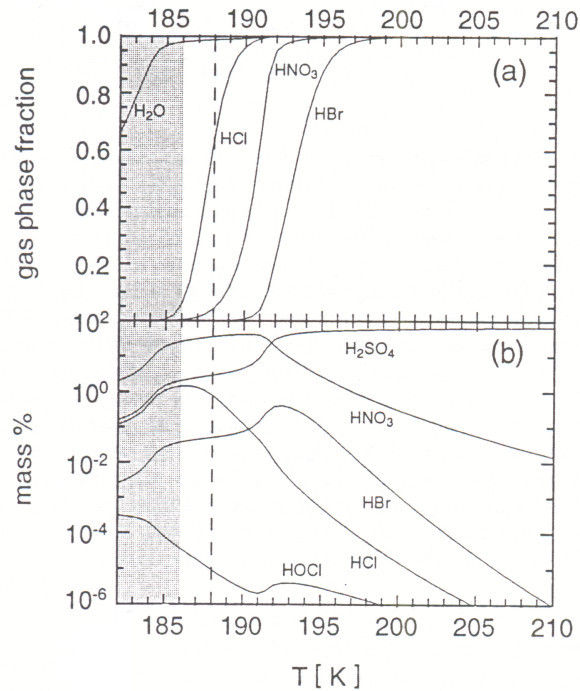


Fig. 6. Partitioning of stratospheric gases into aerosol droplets: (a) fraction of each species in gas phase, and (b) mass fraction of each species in the condensed phase. The vertical dashed line at 188K indicates the ice frost point. The aerosols are unlikely to remain purely liquid more than a couple of degrees below the frost point, so the system properties within the shaded region are hypothetical. The simulated conditions are 50 mbar altitude, with total amounts of each species as follows: H_2O , 5 ppmv; H_2SO_4 , 0.5 ppbv; HNO_3 , 10 ppbv; HCl , 1 ppbv; HBr , 0.01 ppbv; $HOCl$, 0.01 ppbv. The model of Carslaw et al. (1995) was used for these calculations, except for $HOCl$, which Carslaw et al. (1997) calculated using an adapted version of the H_2SO_4 -solution model of Huthwelker et al. (1995). From Carslaw et al. (1997). Copyright 1997 American Geophysical Union. Reproduced by permission of American Geophysical Union.

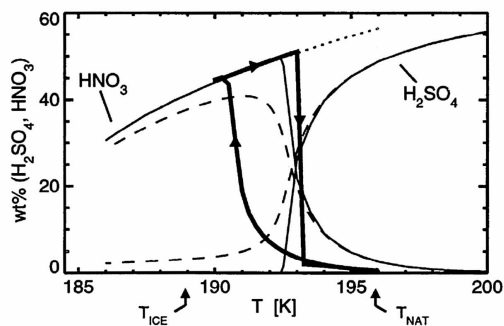


Fig. 7. HNO_3 and H_2SO_4 mass fractions in STS droplets. The total mixing ratios for H_2O and HNO_3 are 5 ppmv and 10 ppbv respectively. The dashed lines represent the changes in particle composition during slow adiabatic cooling of the aerosol. The line solid lines show the changes in particle composition during slow adiabatic cooling without gas-phase depletion, while the dotted line indicates the continuation of the $\text{HNO}_3/\text{H}_2\text{O}$ binary curve to high temperatures. The droplet ensemble is subjected to adiabatic cooling from 196 to 190 K at a rate of 6 K hr^{-1} , followed by adiabatic warming back to 196 K at the same rate. The thick line indicates the evolution of the HNO_3 mass fraction of a droplet, with radii ~ 0.007 and $\sim 0.1 \mu\text{m}$ at temperatures 196 and 190 K respectively, during this event. From Meilinger et al. (1995). Copyright 1995 American Geophysical Union. Reproduced by permission of American Geophysical Union.

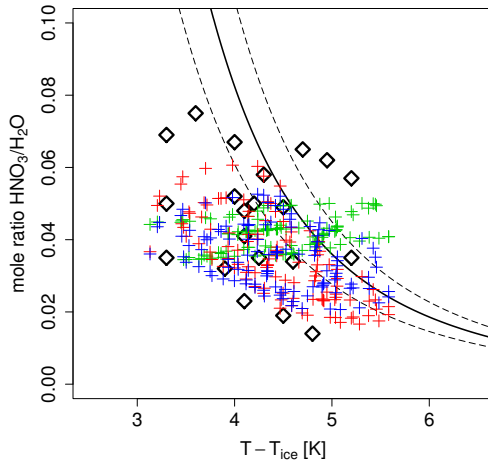


Fig. 8. HNO₃ : H₂O mole ratios in observed STS droplets, plotted as a function of temperature relative to the frost-point T_{ice} . Measurements made by mass spectrometer (diamonds), from Schreiner et al. (1999). Equilibrium mole fractions, calculated after Carslaw et al. (1995), using 5.5 ppmv H₂O, 0.4 ppbv H₂SO₄ and 10 ppbv HNO₃ (solid line). Equilibrium mole fractions for 8 and 12 ppbv HNO₃ are shown as dashed lines. The crosses show the results from a condensed-mass advection model (Lowe, 2003), sampled at fixed radii of 0.09 μm (red), 0.2 μm (blue), and 1.0 μm (green), in a simulation including high frequency temperature fluctuations between 20 and 90 K hr⁻¹. After Voigt et al. (2000b).

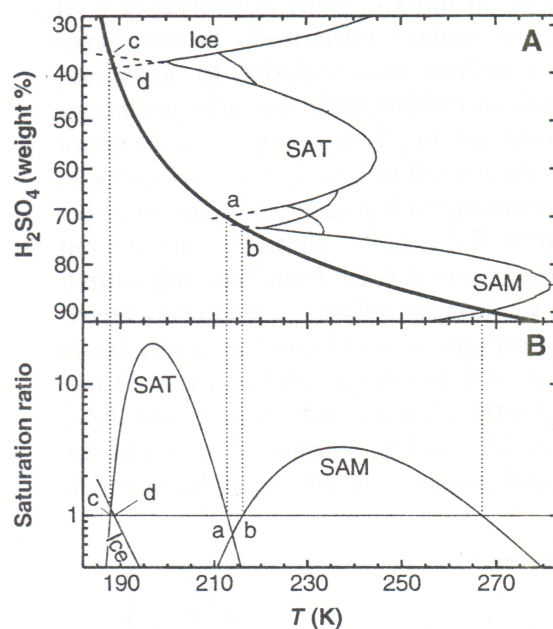


Fig. 9. Comparison of the stability of stratospheric $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ droplets with the different H_2SO_4 solids. The droplets are in equilibrium with a water mixing ratio of 5 ppmv at a 50 mbar altitude. (A) Concentration of droplets (thick solid line) superimposed on the $\text{H}_2\text{SO}_4\text{-H}_2\text{O}$ phase diagram. The melting points of the different solid phases (thin solid lines) are taken from Gable et al. (1950). The major solid phases are water ice (H_2O), SAT ($\text{H}_2\text{SO}_4 \cdot 4\text{H}_2\text{O}$) and SAM ($\text{H}_2\text{SO}_4 \cdot \text{H}_2\text{O}$). The melting points of other minor solid phases ($\text{H}_2\text{SO}_4 \cdot 6.5\text{H}_2\text{O}$, $\cdot 3\text{H}_2\text{O}$, and $\cdot 2\text{H}_2\text{O}$) are also shown. (B) H_2O saturation ratio with respect to ice, SAT, and SAM. Point d is the ice frost point. From Koop and Carslaw (1996). Reprinted with permission from AAAS.

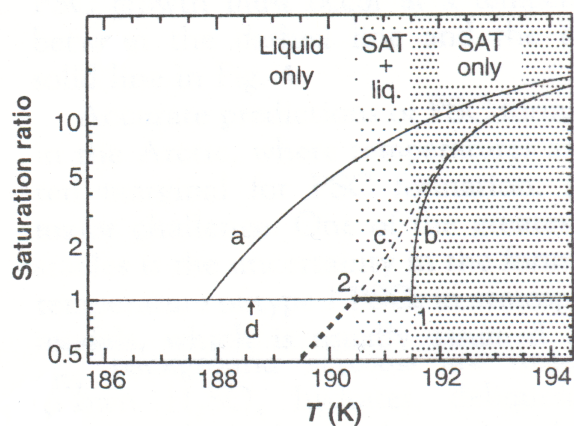


Fig. 10. Saturation ratio of liquid stratospheric aerosols with the same composition as SAT. Line a: pure $\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ droplets (as in Fig. 9). Line b: ternary $\text{HNO}_3/\text{H}_2\text{SO}_4/\text{H}_2\text{O}$ droplets at 50 mbar altitude in equilibrium with 5 ppmv H_2O and a fixed HNO_3 gas phase of 10 ppbv. Line c: same as line b, except with a constant total HNO_3 amount of 10 ppbv, allowing for the partitioning of HNO_3 into the droplets. In the densely shaded region the SAT particles are stable, and therefore remain “dry”. Point 1 is the deliquescence temperature of SAT. Between points 1 and 2 (light shading) a STS film coexists with SAT. At point 2 SAT is completely dissolved; below this temperature only purely liquid droplets exist (thick dashed line). Point d is the ice frost point. All solution ratios calculated using the model of Carslaw et al. (1995). From Koop and Carslaw (1996). Reprinted with permission from AAAS.

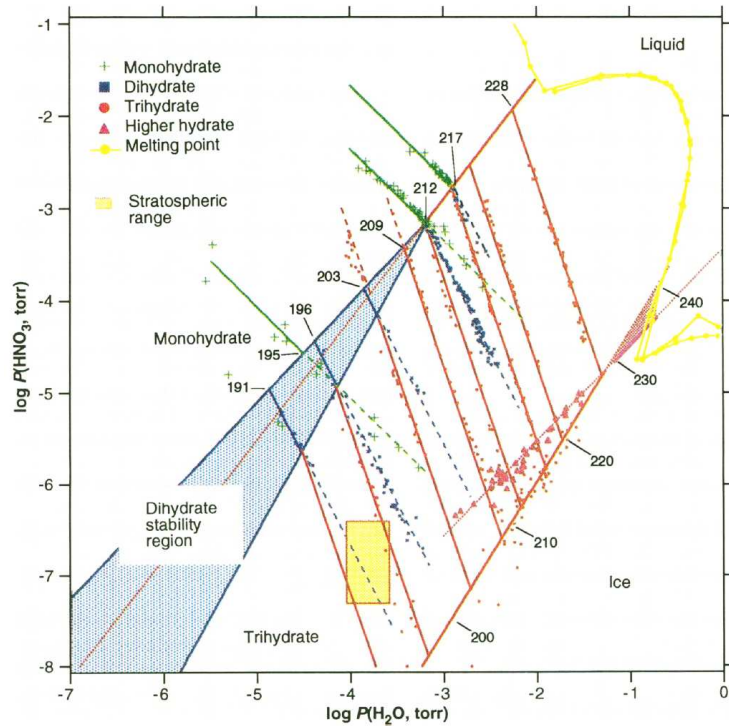


Fig. 11. Nitric acid-water phase diagram (HNO_3 vapour pressure vs. H_2O vapour pressure), with data for the different phases coded by colour and symbol as shown in the key. Lines with positive slopes delineate phase boundaries between (from the lower righthand corner) ice, $\text{HNO}_3 \cdot 3\text{H}_2\text{O}$, $\text{HNO}_3 \cdot 2\text{H}_2\text{O}$, and $\text{HNO}_3 \cdot \text{H}_2\text{O}$ for stable (solid lines) and metastable (dotted lines) phases. The colour-coded lines with negative slopes indicate fits to composition data at the specified temperature. Again solid and dashed lines denote stable and metastable phases, respectively. The blue shaded area indicates the region of stability for $\text{HNO}_3 \cdot 2\text{H}_2\text{O}$. The yellow shaded area indicates the typical range of stratospheric conditions. From Worsnop et al. (1993). Reprinted with permission from AAAS.

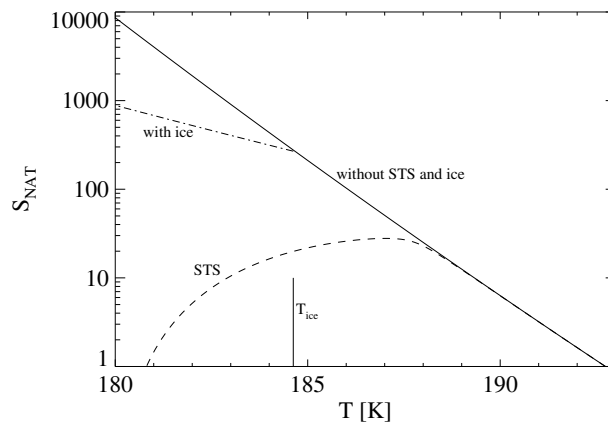


Fig. 12. NAT saturation ratios as a function of temperature, with 5 ppm H_2O , 8 ppb HNO_3 , and 0.2 ppb H_2SO_4 . The solid line indicates the saturation ratio without STS and ice. The dashed line indicates the saturation ratio for a gas-phase in equilibrium with STS. The dashed and dotted line indicates the saturation ratio for a gas-phase in equilibrium with ice. From Luo et al. (2003). Copyright 2003 American Geophysical Union. Reproduced by permission of American Geophysical Union.

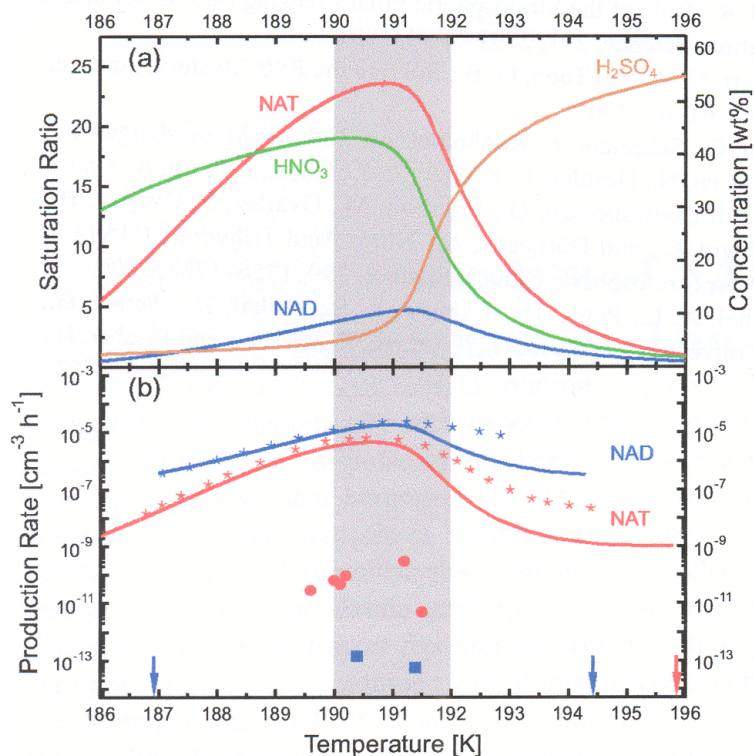


Fig. 13. (a) The compositions (green and orange lines) and the saturation ratios (red and blue lines) of STS aerosols as a function of temperature at 50 mbar with 5 ppmv H₂O, 10 ppbv HNO₃ and 0.5 ppbv H₂SO₄ (Carslaw et al., 1994). The shaded region indicates the temperature range where the saturation ratios over NAD and NAT are at their maximum. (b) Hourly production rates of NAD and NAT particles (blue and red, respectively) per cm³ of air derived from nucleation rate coefficients for the conditions shown in panel (a). The solid lines indicate the production rates in STS droplets calculated using the formulation of Tabazadeh et al. (2001). The squares and circles indicate the production rates derived from the experimental data of Knopf et al. (2002). The stars show values for similar conditions taken from Fig. 1 of Tabazadeh et al. (2001) (produced using a constant total aerosol volume, unlike the other production values which are calculated using temperature dependent aerosol volumes). The arrows indicate the temperatures at which the saturation ratios of NAD (blue) and NAT (red) equal one. From Knopf et al. (2002). By permission of Copernicus Publications on behalf of the European Geosciences Union.

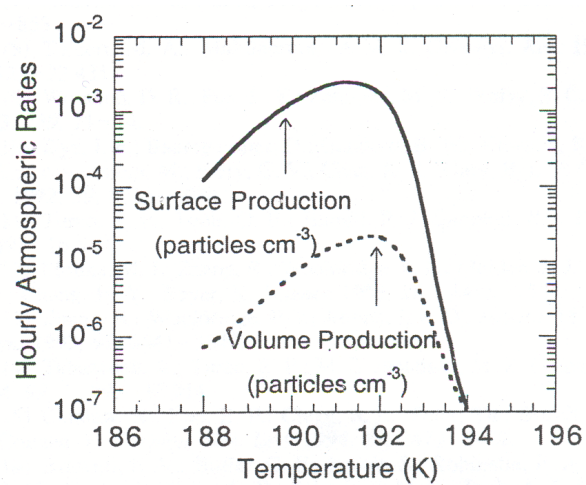


Fig. 14. Comparison of the volume and surface production rates of NAD particles as a function of temperature at 60 mbar with 5 ppmv H_2O , 12 ppbv HNO_3 and 0.5 ppbv H_2SO_4 . Reprinted in part, with permission, from Tabazadeh et al. (2002a) Copyright 2002 American Chemical Society.

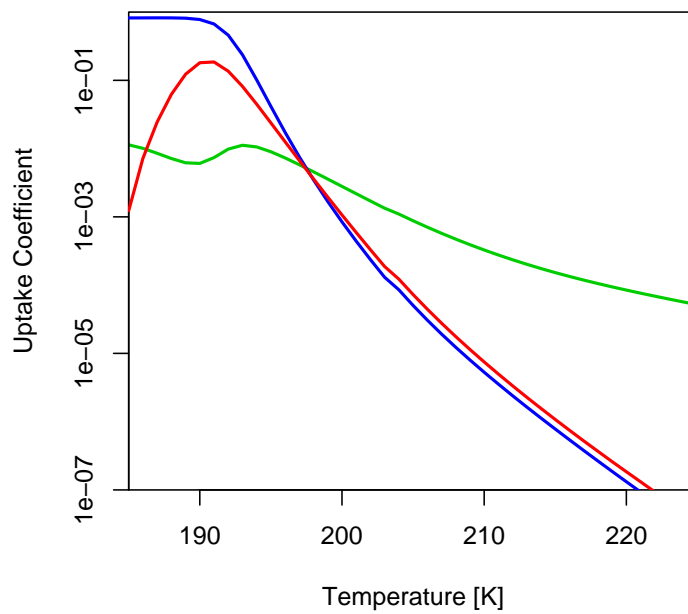


Fig. 15. Reactive uptake coefficients as a function of temperature for heterogeneous reactions occurring in liquid sulphuric acid aerosols. Reaction $\text{ClONO}_2 + \text{H}_2\text{O}$ is represented by the green line, $\text{ClONO}_2 + \text{HCl}$ by the blue line and $\text{HOCl} + \text{HCl}$ by the red line. The uptake coefficients are calculated using the parameterisation of Shi et al. (2001), which takes into account the couplings of these reactions.

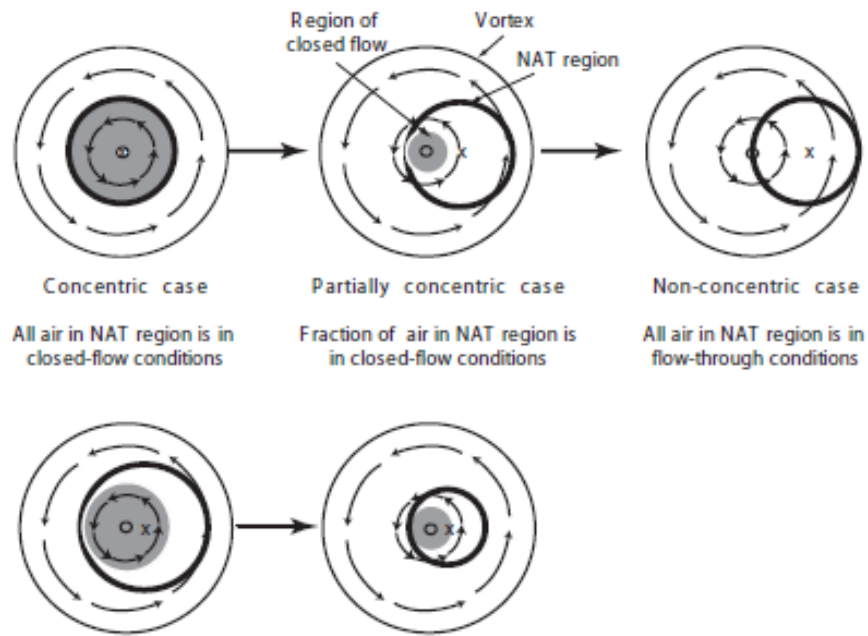


Fig. 16. Schematic of how the concentricity of the vortex and the area of air with temperatures below T_{NAT} (labelled “NAT region”), and the area of the “NAT region”, control the denitrification rate. The “closed-flow” region is shaded. The upper row of diagrams illustrate the effect of shifting the NAT region away from the centre of the vortex; the lower row the effect of reducing the size of the “NAT region.” From Mann et al. (2003). By permission of Copernicus Publications on behalf of the European Geosciences Union.

PSC Type	Backscatter ratio	α	Depolarisation, %
1a	0.2–0.5	0.4	30–50
1b	2–7	2–3	0.5–2.5
2	> 10	< 0.8	> 10

Table 1

Typical PSC lidar properties when measured using a wavelength of $0.603 \mu\text{m}$ (Browell et al., 1990).